

Is the 2MASS dipole convergent?

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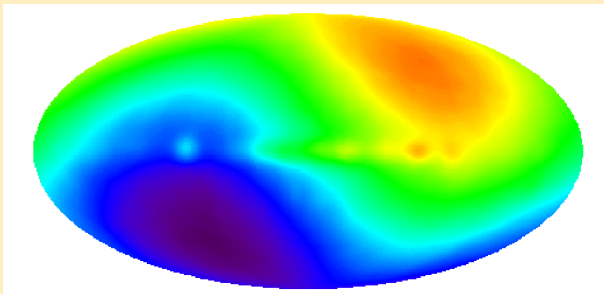
The β parameter

- Comparison of the Local Group (LG) velocity with its acceleration allows, within linear theory, to calculate $\beta \equiv \Omega_m^{0.55}/b$:

$$v_{\text{LG}} \propto \beta g_{\text{LG}}$$

- The velocity of the LG known from the CMB dipole,

$$v_{\text{LG}} = 622 \pm 35 \text{ km/s towards } (l, b) = (272 \pm 3^\circ, 28 \pm 5^\circ)$$



COBE, ©NASA

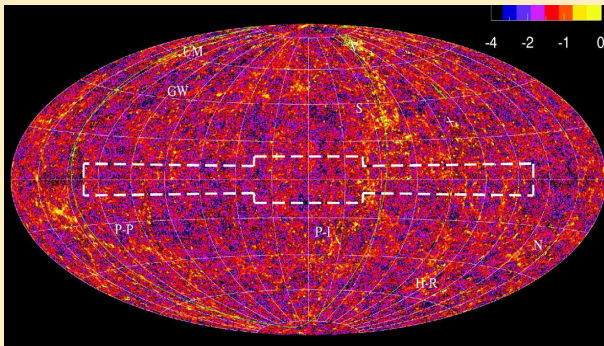


Dipole of the 2MASS survey

- The acceleration of the LG ('clustering dipole') is estimated from the *2-Micron All-Sky Survey* (**2MASS**):

$$\mathbf{g}_{\text{LG}} \sim \rho_{\text{L}}^{-1} \sum_i S_i \hat{\mathbf{r}}_i$$

where S_i – observed flux of the i -th galaxy in the K_S band

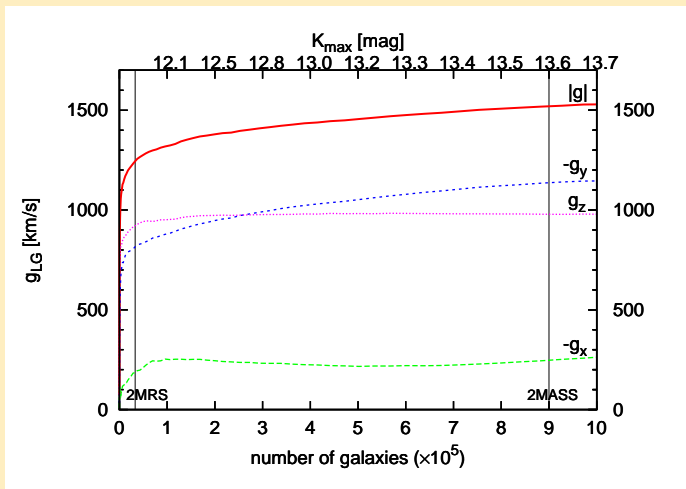


Maller *et al.* 2003, ApJL



Is the 2MASS dipole convergent?

Growth of the 2MASS dipole (1)

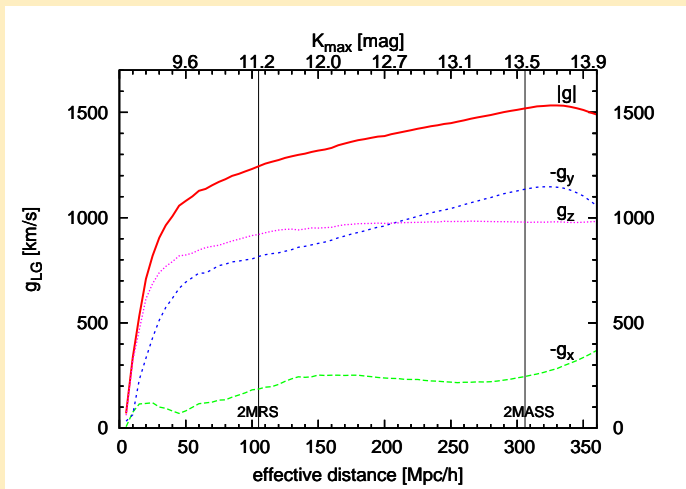


(1): as a function of *number of galaxies*



Is the 2MASS dipole convergent?

Growth of the 2MASS dipole (2)

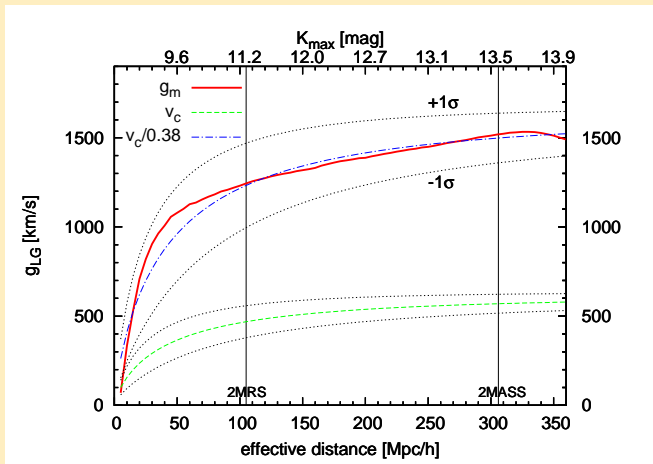


(2): as a function of *effective distance*



Is this expected?

Comparison with theoretical expectations

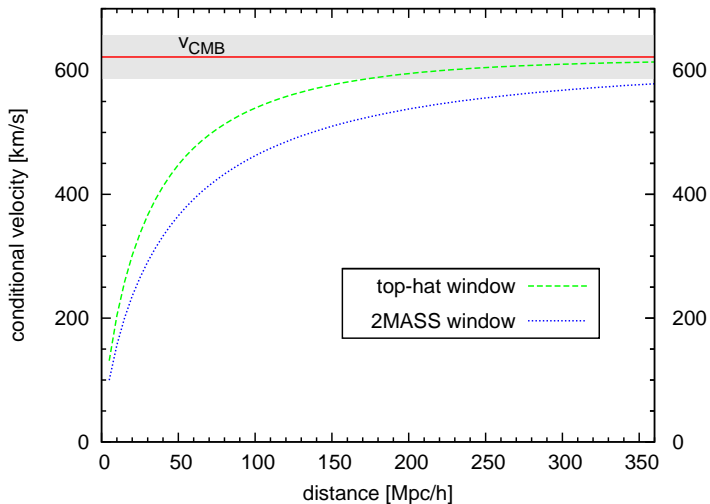


Conditional mean velocity, v_c (JVV 1990; KL 1990), for Λ -CDM power spectrum and cosmological parameters from 5-year WMAP



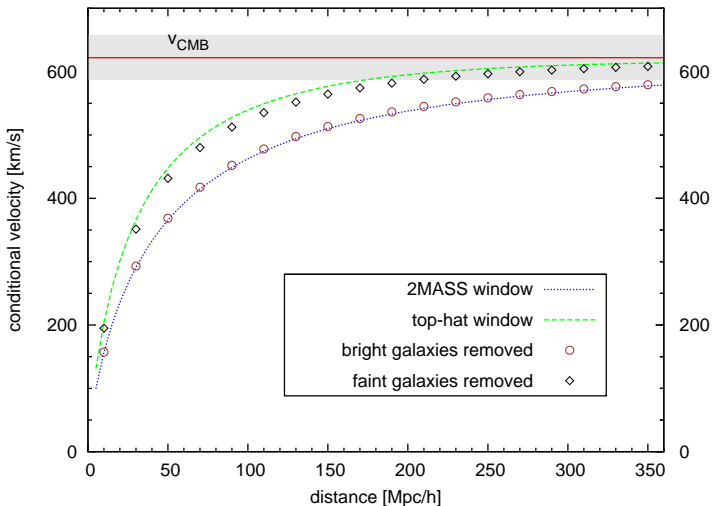
Theoretical expectations

Effect of different observational windows



Theoretical expectations

Whence the difference?



Growth of the 2MASS dipole: results

Summary



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- The dipole does **not** converge up to about $300 \text{ Mpc}/h$
- Inconsistent with theoretically predicted convergence of the dipole for about $200 \text{ Mpc}/h$ in ΛCDM for the **top-hat** observational window



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- However, observed growth of the dipole as expected in ΛCDM for the **2MASS** observational window
- A preliminary value of $\beta \equiv \Omega_m^{0.55} / b$ is about 0.38
- Combined with linear biasing from Maller et al. (2005), $b_K \simeq 1.1 \pm 0.2$, it gives

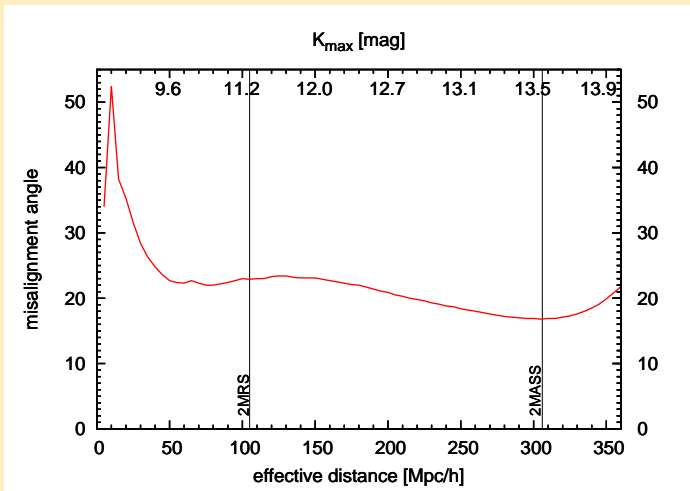
$$\Omega_m \simeq 0.21 \pm 0.07$$



Thank you



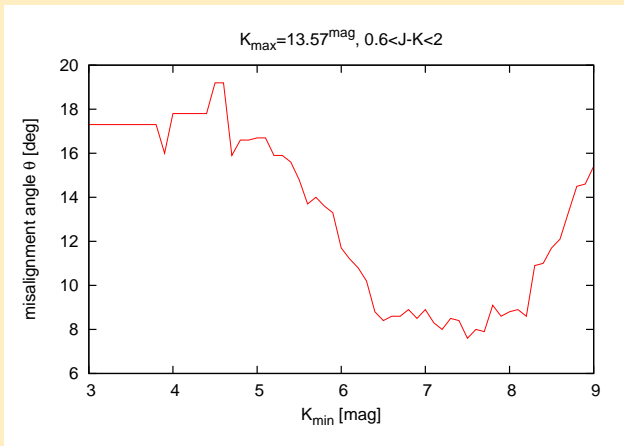
Misalignment angle (1)



(1): as a function of *effective distance*



Misalignment angle (2)

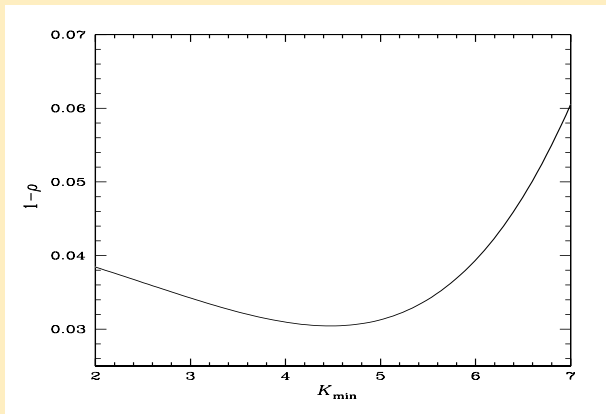


(2): as a function of *minimal* magnitude of *included* galaxies

(for $K_{\max} = 13.57$ mag corresponding to $R_{\text{eff}} \simeq 300$ Mpc/h)



Misalignment angle: theory

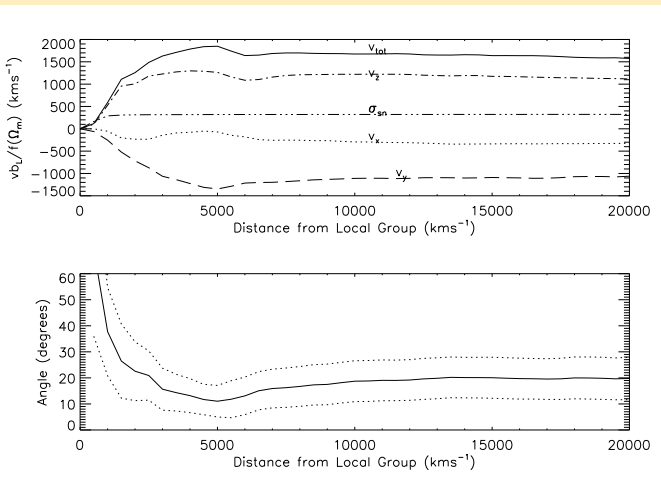
Chodorowski *et al.* 2007, MNRAS

The misalignment angle $\theta \propto \sqrt{1-\rho}$, where ρ is the cross-correlation coefficient of \mathbf{v}_{LG} and \mathbf{g}_{LG} . In above figure, decorrelation only due to nonlinear effects.



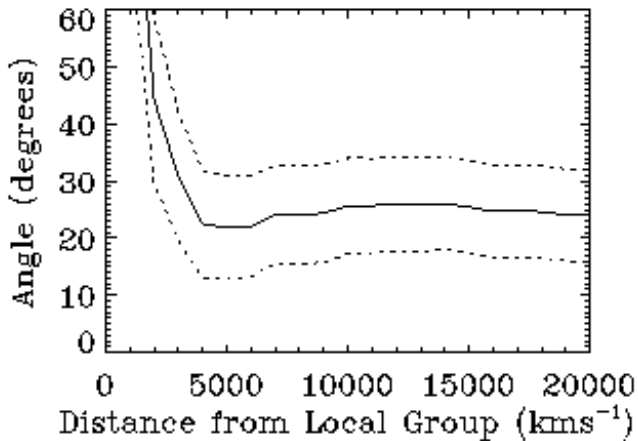
Erdođdu et al. 2006, arXiv

2MASS Redshift Survey

For 2MRS, $K_{\text{max}} = 11.25 \text{ mag}$, corresponding to $R_{\text{eff}} \simeq 100 \text{ Mpc}/h$ 

Erdođdu et al. 2006, MNRAS

2MASS Redshift Survey



CMB rest frame

