

XXXIXth Rencontres de Moriond: QCD and High-Energy Hadronic Interactions
La Thuile, March 28th-April 4th 2004

Glueon radiation off massive quarks in a QCD medium

Néstor Armesto, Carlos A. Salgado and Urs Achim Wiedemann

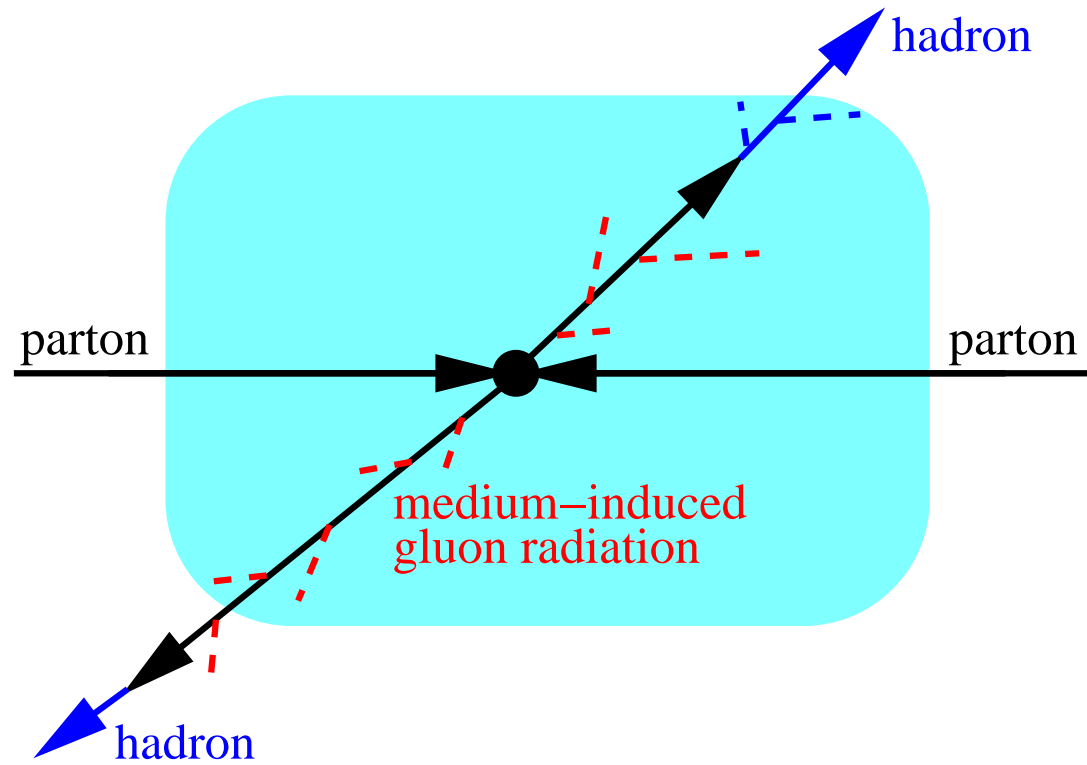
Department of Physics, Theory Division, CERN

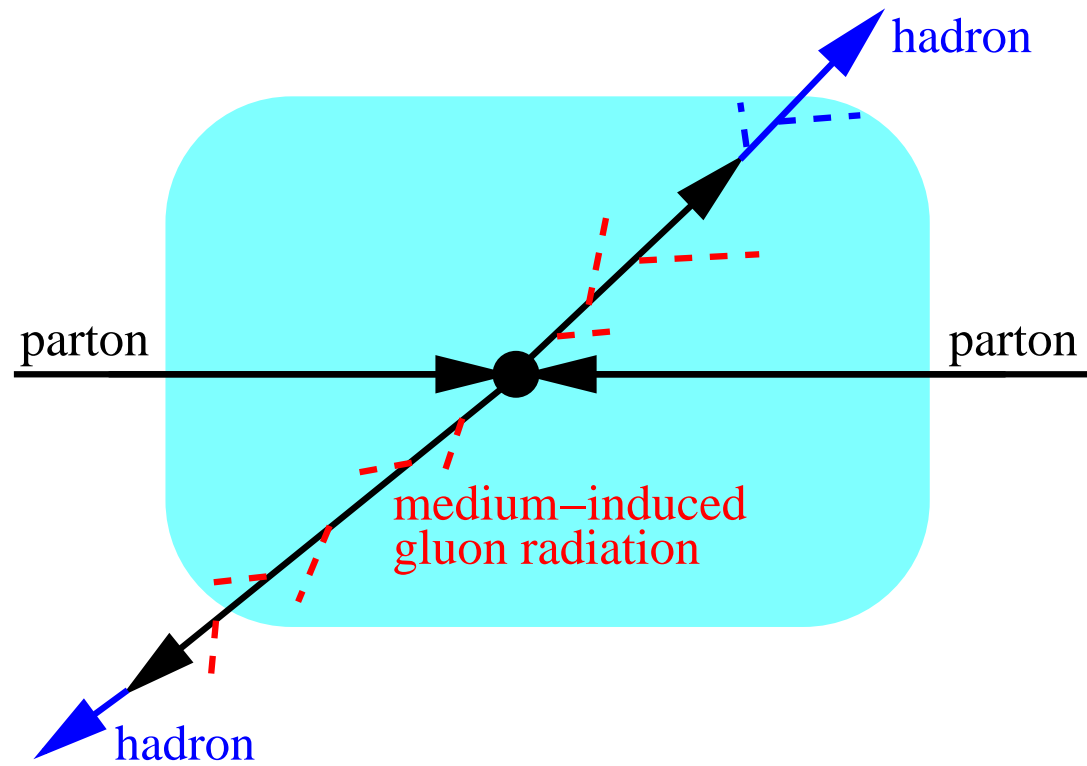
1. Introduction.
2. Results.
3. Conclusions.

Preprint CERN-TH/2003-301 (hep-ph/0312106), PRD to appear.

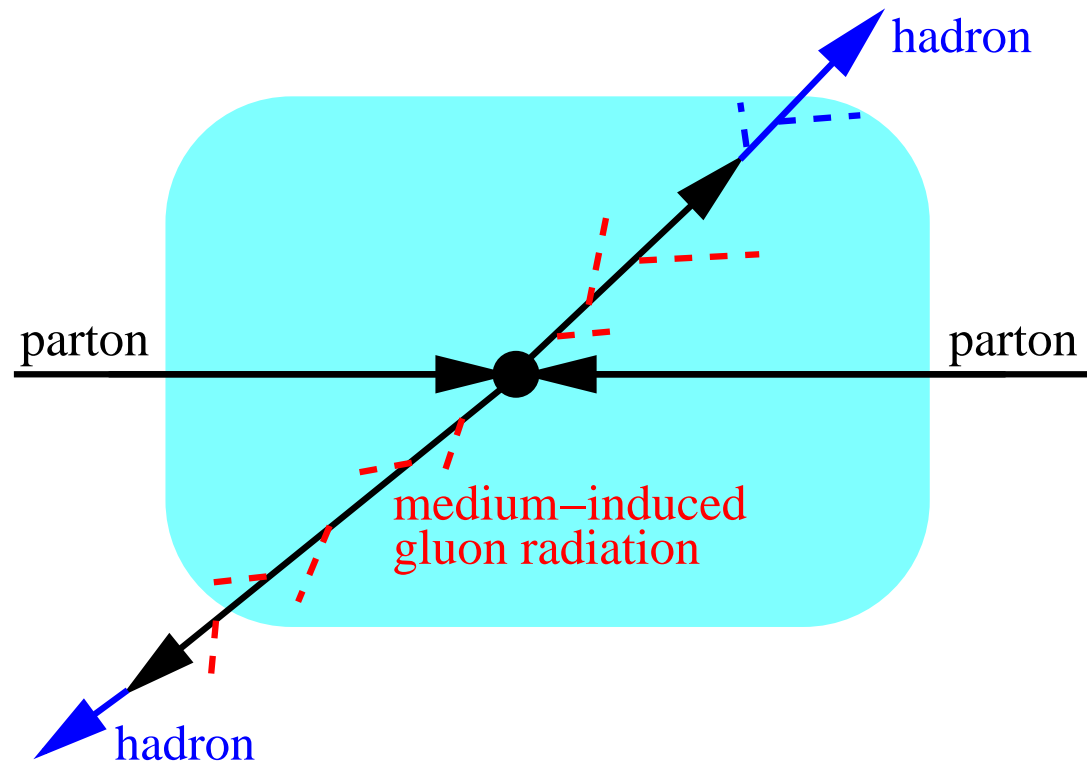
Related talks: Dainese, d'Enterria, Ferreiro, Kharzeev, Liu and Wang.

1. Introduction:



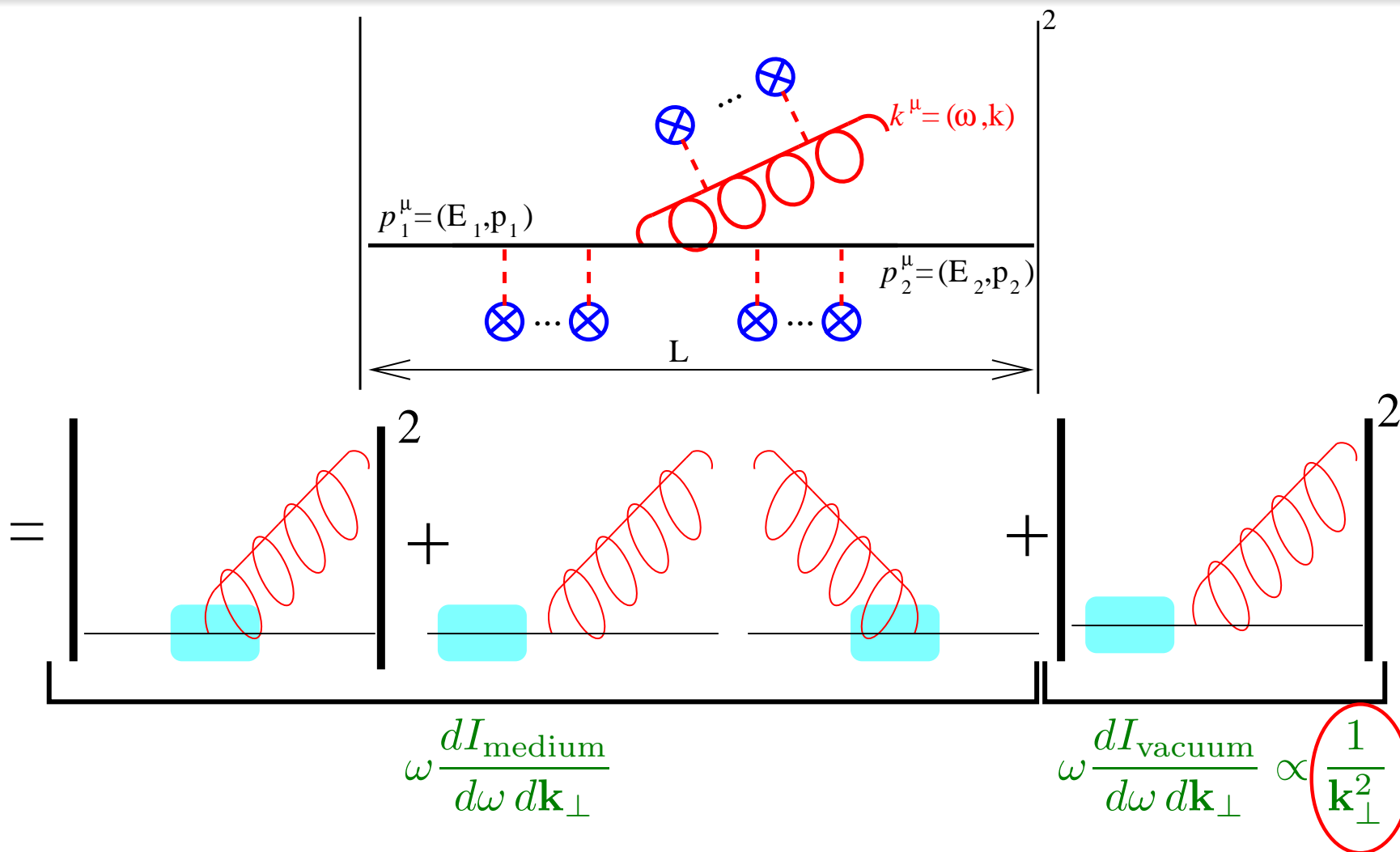


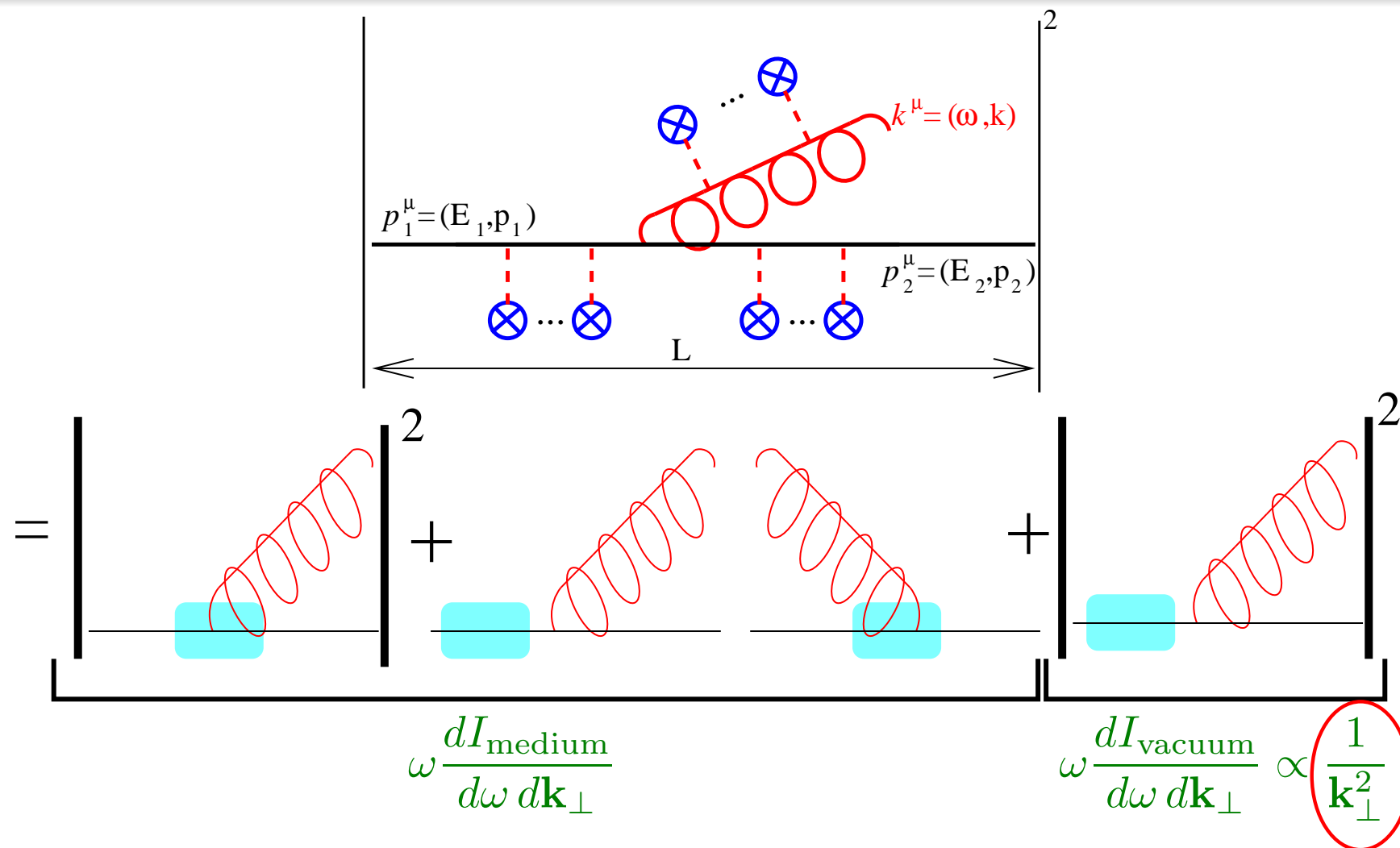
- **Medium-induced gluon radiation implies:**
 - ⇒ Energy degradation of the leading parton.
 - ⇒ Broadening of the parton shower.
 - ⇒ Increase of the associated hadron multiplicity.



- **Medium-induced gluon radiation implies:**
 - ⇒ Energy degradation of the leading parton.
 - ⇒ Broadening of the parton shower.
 - ⇒ Increase of the associated hadron multiplicity.
- ♣ **How does the medium-induced gluon radiation off a massive quark differ from that off a massless parton?**

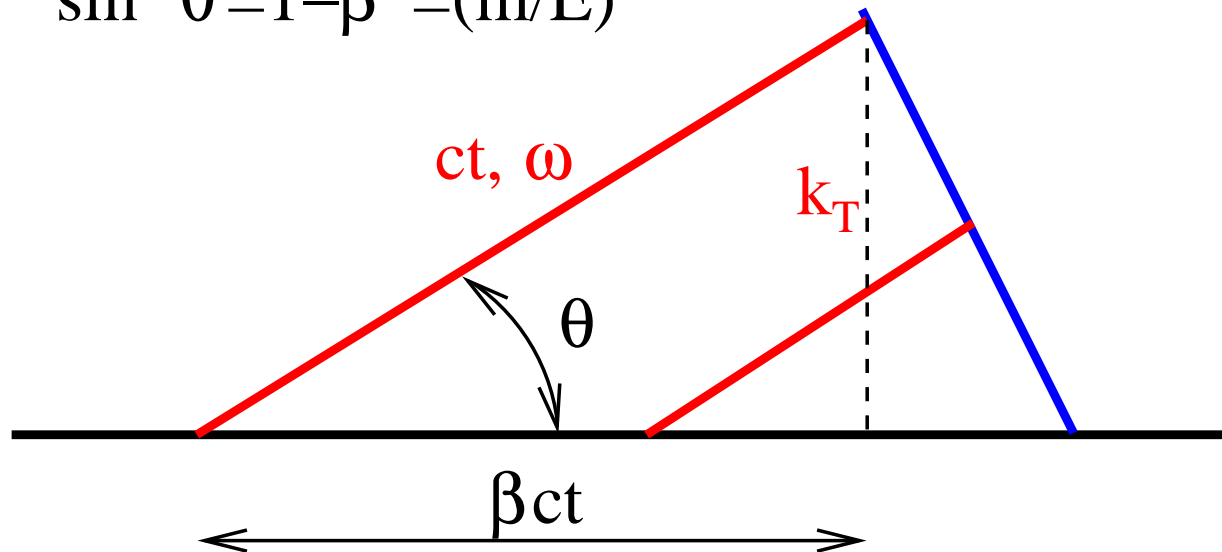
Medium-induced gluon radiation off massless quarks:





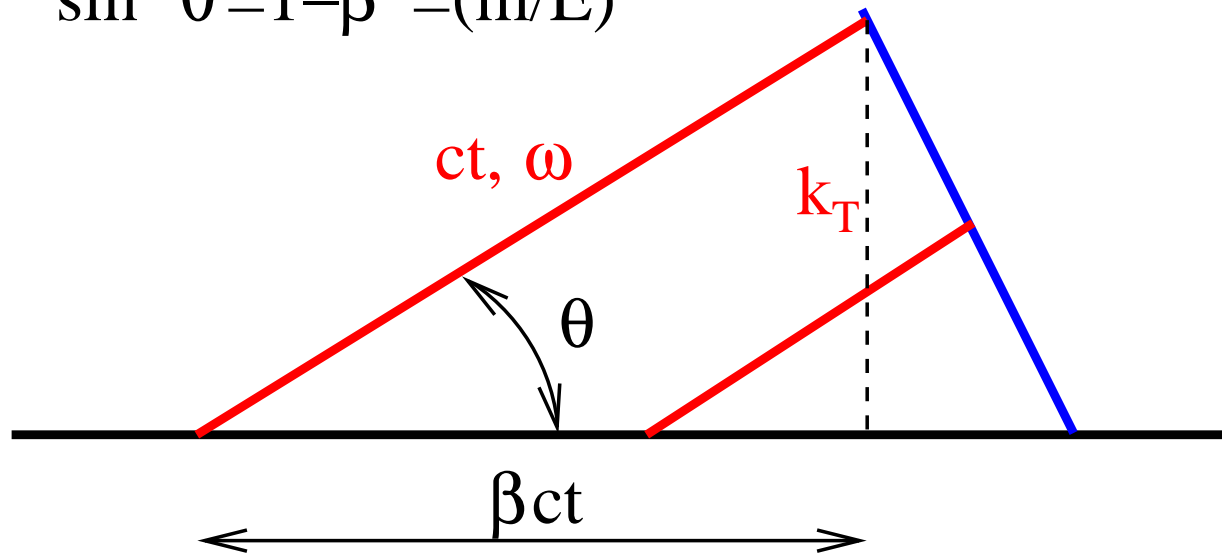
- The BDMPS formalism describes the medium-modification of the vacuum radiation pattern (Baier-Dokshitzer-Mueller-Peigné-Schiff; Zakharov; Wiedemann; Gyulassy-Levai-Vitev; Wang-Guo; see reviews in: Baier, Schiff, Zakharov, ARNPS50(00)37; Kovner, Wiedemann in Quark Gluon Plasma 3; Gyulassy, Vitev, Wang, Zhang, ibid).

$$\sin^2 \theta = 1 - \beta^2 = (m/E)^2$$



- Gluon radiation **in the vacuum** is modified by a mass of the parent quark: radiation for angles $\theta < m/E$ is suppressed, the **dead cone effect** (Dokshitzer, Khoze, Troyan, JPG17(91)1481):

$$\sin^2 \theta = 1 - \beta^2 = (m/E)^2$$

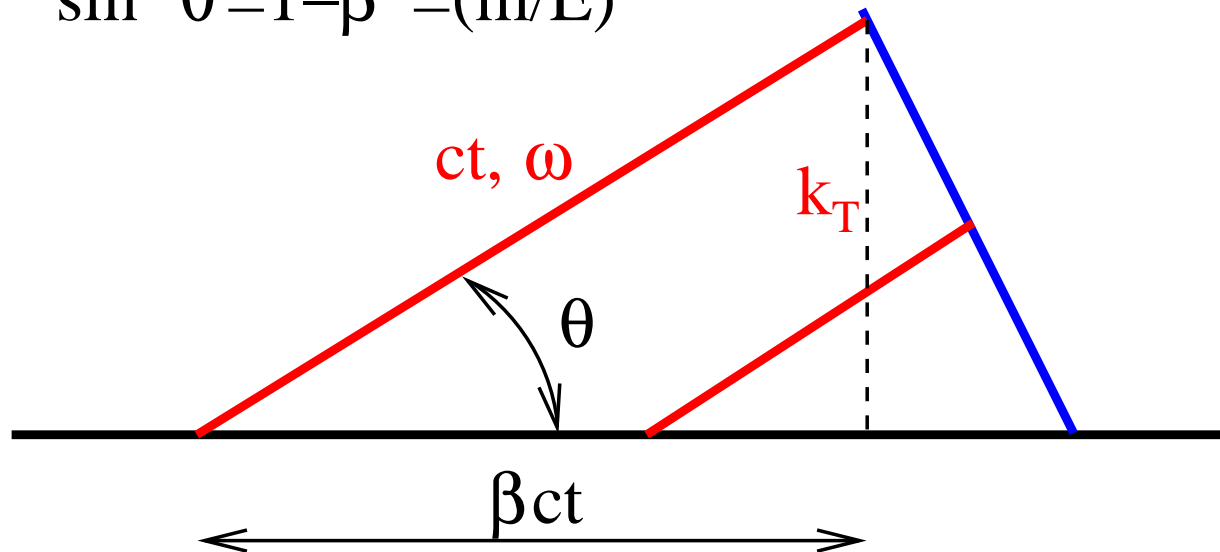


- Gluon radiation **in the vacuum** is modified by a mass of the parent quark: radiation for angles $\theta < m/E$ is suppressed, the **dead cone effect** (Dokshitzer, Khoze, Troyan, JPG17(91)1481):

$$\frac{1}{\mathbf{k}_{\perp}^2} \rightarrow \frac{1}{\mathbf{k}_{\perp}^2} \left[\frac{\mathbf{k}_{\perp}^2}{\mathbf{k}_{\perp}^2 + \left(\frac{m\omega}{E}\right)^2} \right]^2 \equiv \frac{1}{\mathbf{k}_{\perp}^2} F\left(\mathbf{k}_{\perp}^2, \frac{m\omega}{E}\right).$$

Dead cone factor

$$\sin^2 \theta = 1 - \beta^2 = (m/E)^2$$

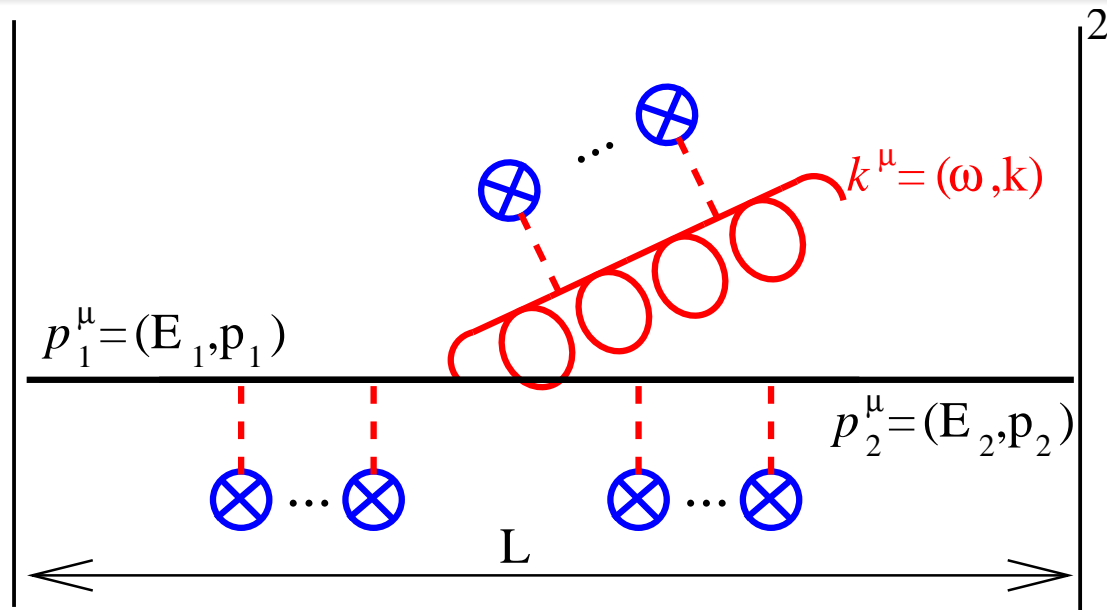


- Gluon radiation **in the vacuum** is modified by a mass of the parent quark: radiation for angles $\theta < m/E$ is suppressed, the **dead cone effect** (Dokshitzer, Khoze, Troyan, JPG17(91)1481):

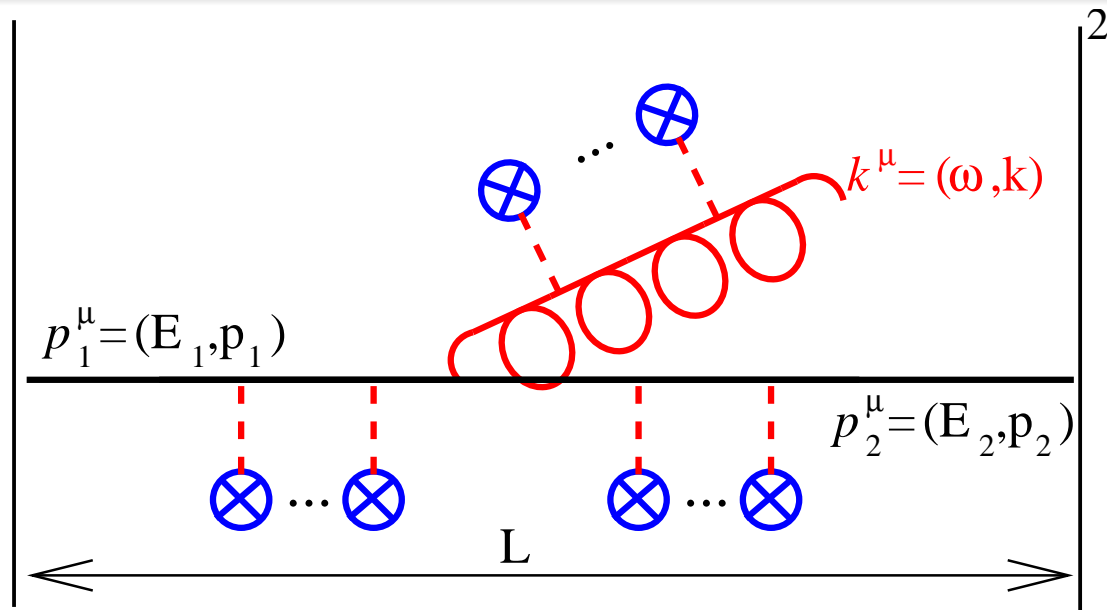
$$\frac{1}{\mathbf{k}_{\perp}^2} \rightarrow \frac{1}{\mathbf{k}_{\perp}^2} \left[\frac{\mathbf{k}_{\perp}^2}{\mathbf{k}_{\perp}^2 + \left(\frac{m\omega}{E}\right)^2} \right]^2 \equiv \frac{1}{\mathbf{k}_{\perp}^2} \underbrace{F\left(\mathbf{k}_{\perp}^2, \frac{m\omega}{E}\right)}_{\text{Dead cone factor}}.$$

- Dokshitzer and Kharzeev (PLB519(01)199) proposed that **medium-induced gluon radiation is reduced by the same effect**. In this first exploratory study:

$$\omega \frac{dI_{\text{medium}}^{m>0}}{d\omega} = \omega \frac{dI_{\text{medium}}^{m=0}}{d\omega} F\left(\langle \mathbf{k}_{\perp}^2 \rangle, \frac{m\omega}{E}\right).$$

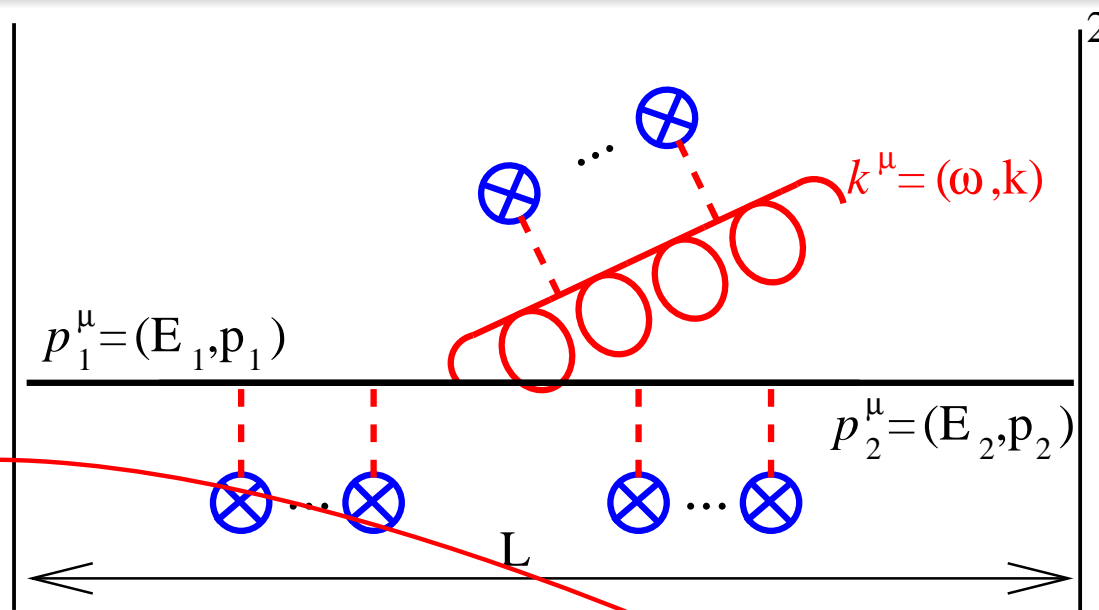


$$\omega \frac{dI}{d\omega d\mathbf{k}_\perp} = \frac{\alpha_s C_F}{(2\pi)^2 \omega^2} 2\text{Re} \int_0^\infty dy_l \int_{y_l}^\infty d\bar{y}_l e^{i\bar{q}(y_l - \bar{y}_l)} \int d\mathbf{u} e^{-i\mathbf{k}_\perp \cdot \mathbf{u}} \times e^{-\frac{1}{2} \int_{\bar{y}_l}^\infty d\xi n(\xi) \sigma(\mathbf{u})} \frac{\partial}{\partial \mathbf{y}} \cdot \frac{\partial}{\partial \mathbf{u}} \int_{\mathbf{y}=0=\mathbf{r}(y_l)}^{\mathbf{u}=\mathbf{r}(\bar{y}_l)} \mathcal{D}\mathbf{r} \exp \left[i \int_{y_l}^{\bar{y}_l} d\xi \frac{\omega}{2} \left(\dot{\mathbf{r}}^2 - \frac{n(\xi) \sigma(\mathbf{r})}{i\omega} \right) \right].$$



$$\omega \frac{dI}{d\omega d\mathbf{k}_\perp} = \frac{\alpha_s C_F}{(2\pi)^2 \omega^2} 2\text{Re} \int_0^\infty dy_l \int_{y_l}^\infty d\bar{y}_l e^{i\bar{q}(y_l - \bar{y}_l)} \int d\mathbf{u} e^{-i\mathbf{k}_\perp \cdot \mathbf{u}} \times e^{-\frac{1}{2} \int_{\bar{y}_l}^\infty d\xi n(\xi) \sigma(\mathbf{u})} \frac{\partial}{\partial \mathbf{y}} \cdot \frac{\partial}{\partial \mathbf{u}} \int_{\mathbf{y}=0=\mathbf{r}(y_l)}^{\mathbf{u}=\mathbf{r}(\bar{y}_l)} \mathcal{D}\mathbf{r} \exp \left[i \int_{y_l}^{\bar{y}_l} d\xi \frac{\omega}{2} \left(\dot{\mathbf{r}}^2 - \frac{n(\xi) \sigma(\mathbf{r})}{i\omega} \right) \right].$$

- Information about the medium contained in $n(\xi) \sigma(\mathbf{r}) \simeq \frac{1}{2} \hat{q}(\xi) \mathbf{r}^2$.
The only model parameter is $\hat{q}(\xi) \simeq \langle \mathbf{q}_\perp^2 \rangle_{\text{medium}} / \lambda_{\text{mfp}}$.



$$\omega \frac{dI}{d\omega d\mathbf{k}_\perp} = \frac{\alpha_s C_F}{(2\pi)^2 \omega^2} 2\text{Re} \int_0^\infty dy_l \int_{y_l}^\infty d\bar{y}_l (e^{i\bar{q}(y_l - \bar{y}_l)}) \int d\mathbf{u} e^{-i\mathbf{k}_\perp \cdot \mathbf{u}} \times e^{-\frac{1}{2} \int_{\bar{y}_l}^\infty d\xi n(\xi) \sigma(\mathbf{u})} \frac{\partial}{\partial \mathbf{y}} \cdot \frac{\partial}{\partial \mathbf{u}} \int_{\mathbf{y}=0=\mathbf{r}(y_l)}^{\mathbf{u}=\mathbf{r}(\bar{y}_l)} D\mathbf{r} \exp \left[i \int_{y_l}^{\bar{y}_l} d\xi \frac{\omega}{2} \left(\dot{\mathbf{r}}^2 - \frac{n(\xi) \sigma(\mathbf{r})}{i\omega} \right) \right].$$

- Information about the medium contained in $n(\xi) \sigma(\mathbf{r}) \simeq \frac{1}{2} \hat{q}(\xi) \mathbf{r}^2$.
The only model parameter is $\hat{q}(\xi) \simeq \langle \mathbf{q}_\perp^2 \rangle_{\text{medium}} / \lambda_{\text{mfp}}$.
- All mass effects: $\bar{q} = p_1 - p_2 - k \simeq \frac{x^2 m^2}{2\omega}$, $x = \omega/E \ll 1$.

1) Vacuum term: dead cone,

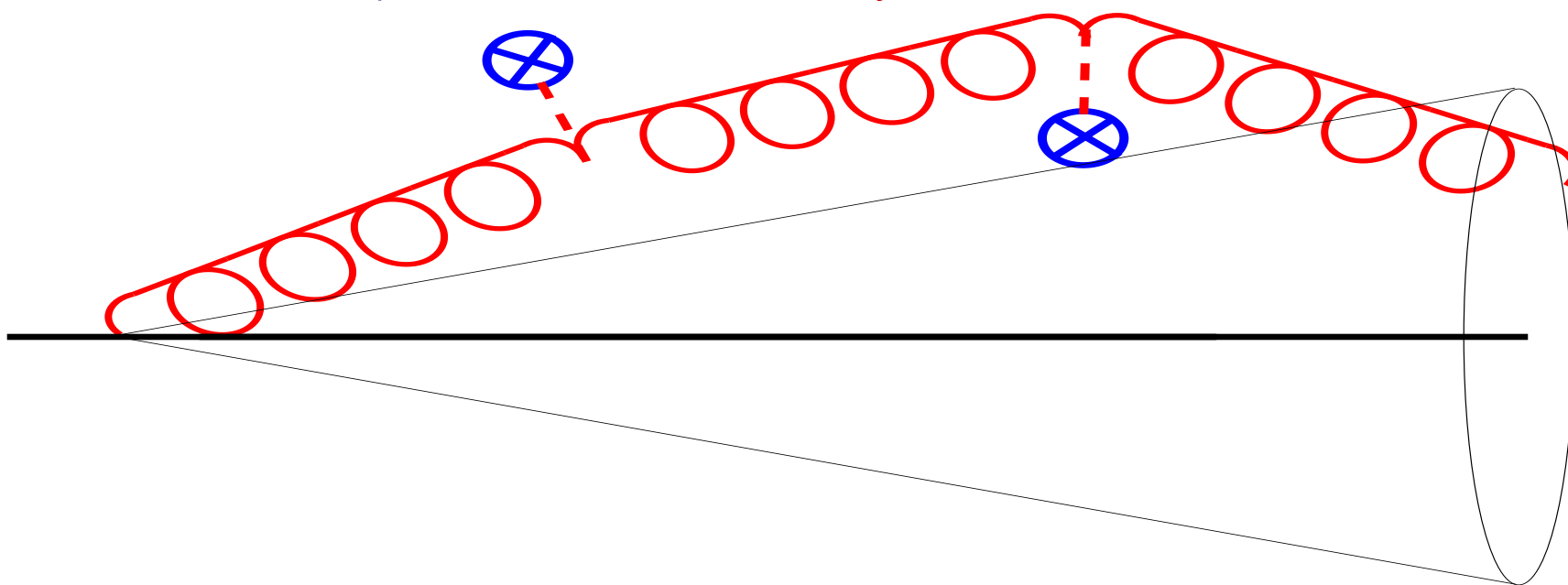
$$\omega \frac{dI_{\text{vacuum}}}{d\omega d\mathbf{k}_{\perp}} \propto \frac{1}{\mathbf{k}_{\perp}^2} F\left(\mathbf{k}_{\perp}^2, \frac{m\omega}{E}\right) = \frac{1}{\mathbf{k}_{\perp}^2} \left[\frac{\mathbf{k}_{\perp}^2}{\mathbf{k}_{\perp}^2 + \left(\frac{m\omega}{E}\right)^2} \right]^2.$$

1) Vacuum term: dead cone,

$$\omega \frac{dI_{\text{vacuum}}}{d\omega d\mathbf{k}_{\perp}} \propto \frac{1}{\mathbf{k}_{\perp}^2} F\left(\mathbf{k}_{\perp}^2, \frac{m\omega}{E}\right) = \frac{1}{\mathbf{k}_{\perp}^2} \left[\frac{\mathbf{k}_{\perp}^2}{\mathbf{k}_{\perp}^2 + \left(\frac{m\omega}{E}\right)^2} \right]^2.$$

2) Medium term:

Naive: gluon moves into the dead cone due to multiple scattering (Brownian motion) \Rightarrow the dead cone may be filled.

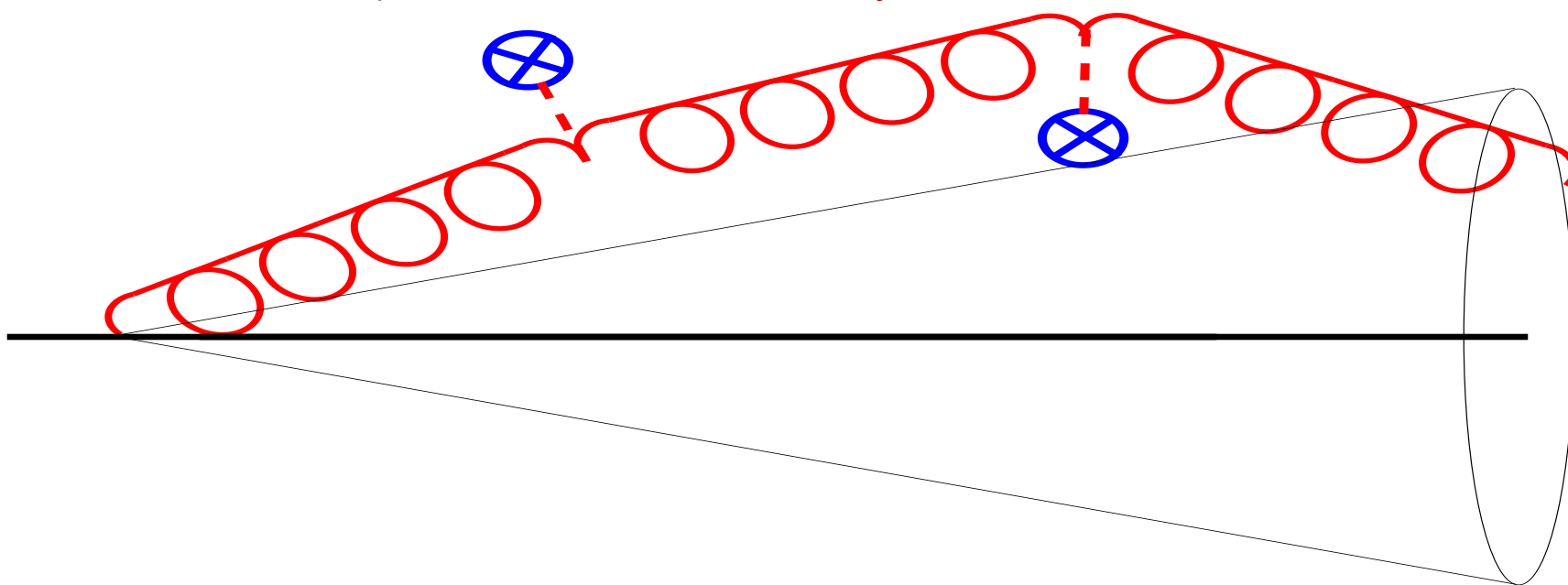


1) Vacuum term: dead cone,

$$\omega \frac{dI_{\text{vacuum}}}{d\omega d\mathbf{k}_{\perp}} \propto \frac{1}{\mathbf{k}_{\perp}^2} F\left(\mathbf{k}_{\perp}^2, \frac{m\omega}{E}\right) = \frac{1}{\mathbf{k}_{\perp}^2} \left[\frac{\mathbf{k}_{\perp}^2}{\mathbf{k}_{\perp}^2 + \left(\frac{m\omega}{E}\right)^2} \right]^2.$$

2) Medium term:

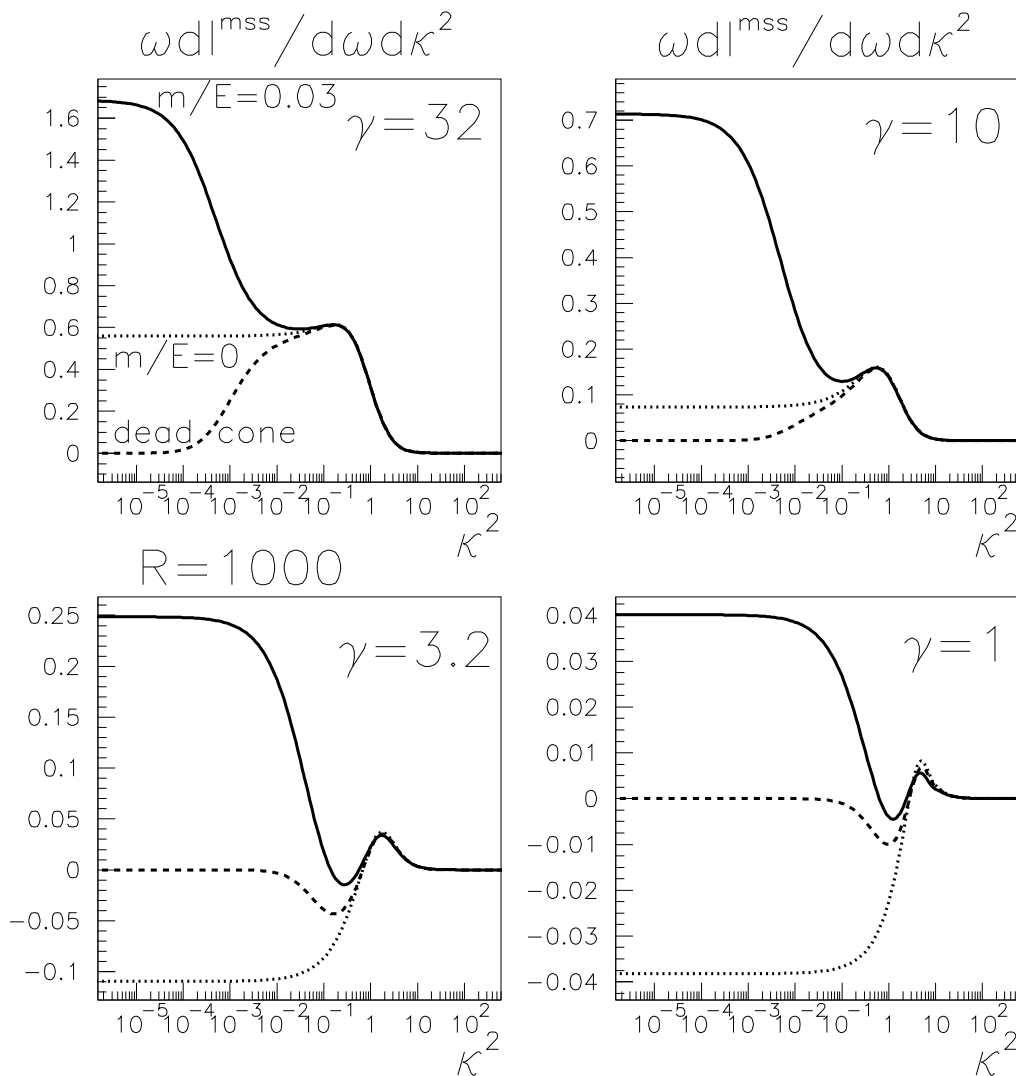
Naive: gluon moves into the dead cone due to multiple scattering (Brownian motion) \Rightarrow the dead cone may be filled.



Technically: competition between interference and rescattering \Rightarrow look at the numerical results.

2. Results (I) – the dead cone is filled:

Néstor Armesto

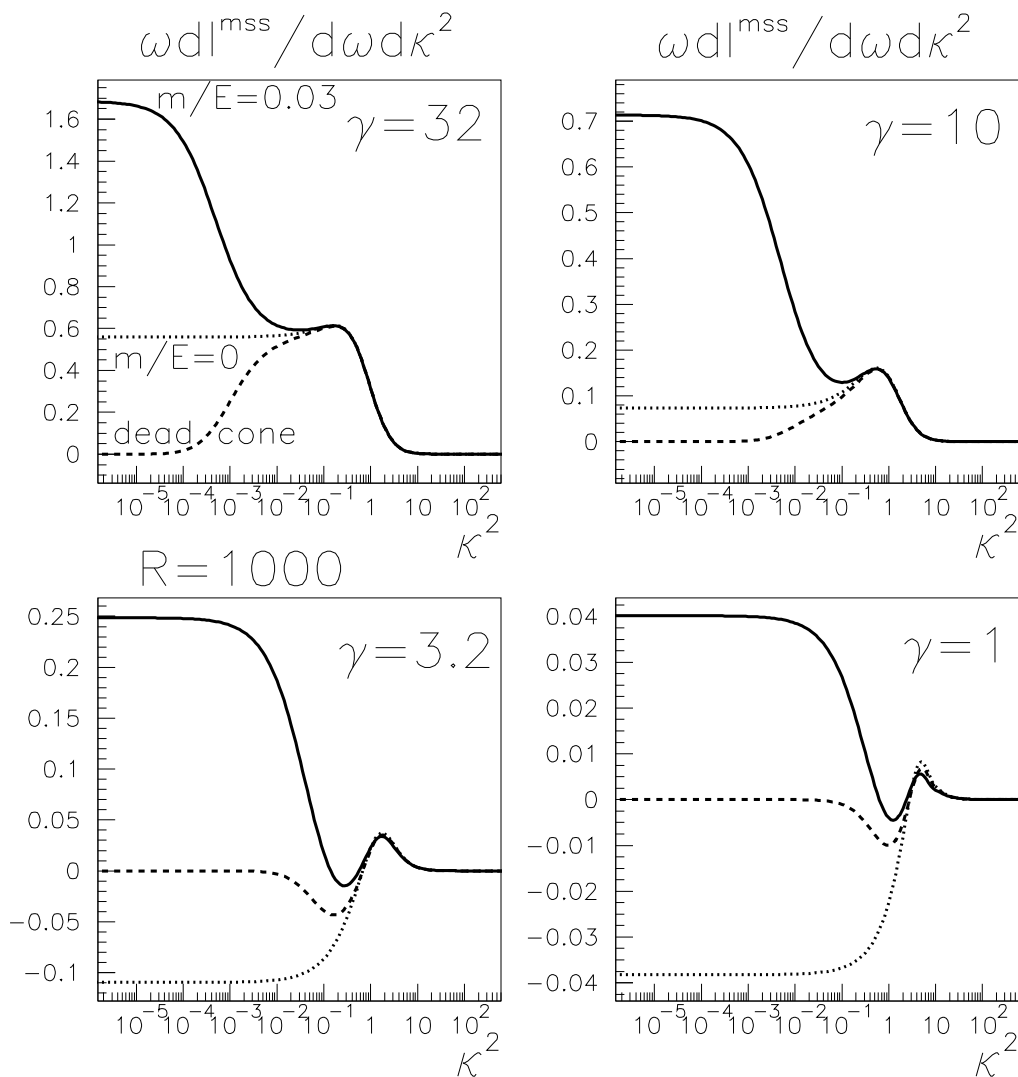


- $\kappa^2 = \frac{k_{\perp}^2}{\hat{q}L}$, $\omega_c = \hat{q}L^2/2$,
 $R = \omega_c L$, $\gamma = \omega_c/\omega$,
 $M^2 = \frac{x^2 m^2}{\hat{q}L}$.

- Dead cone \equiv massless
 $\times F(\kappa, M) = \left(\frac{\kappa^2}{\kappa^2 + M^2} \right)^2$.

2. Results (I) – the dead cone is filled:

Néstor Armesto

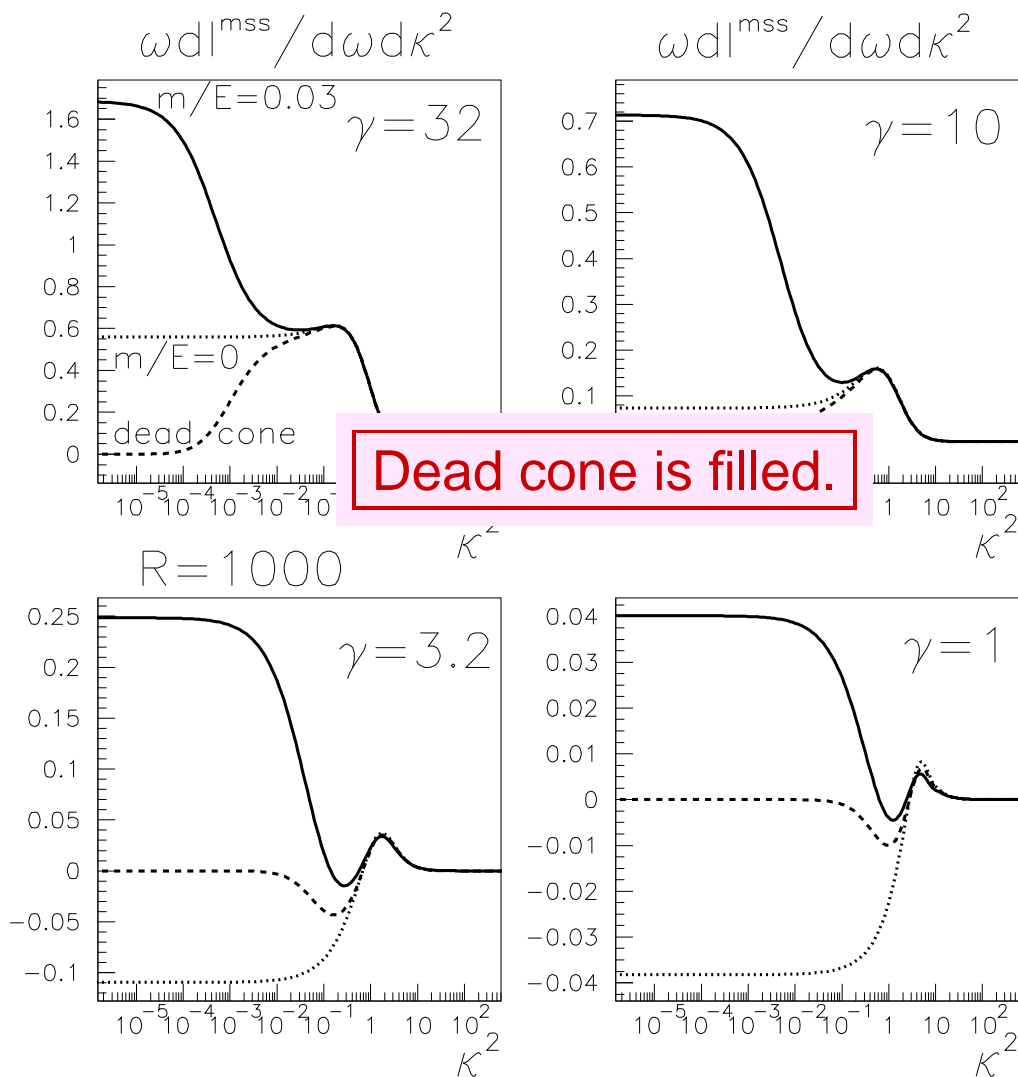


- $\kappa^2 = \frac{k_{\perp}^2}{\hat{q}L}$, $\omega_c = \hat{q}L^2/2$,
 $R = \omega_c L$, $\gamma = \omega_c/\omega$,
 $M^2 = \frac{x^2 m^2}{\hat{q}L}$.

- Dead cone \equiv massless
 $\times F(\kappa, M) = \left(\frac{\kappa^2}{\kappa^2 + M^2} \right)^2$.
- $I_{\text{med}} + I_{\text{vac}} \geq 0$.
- For $m > 0$, $I_{\text{med}} \geq 0$.

2. Results (I) – the dead cone is filled:

Néstor Armesto

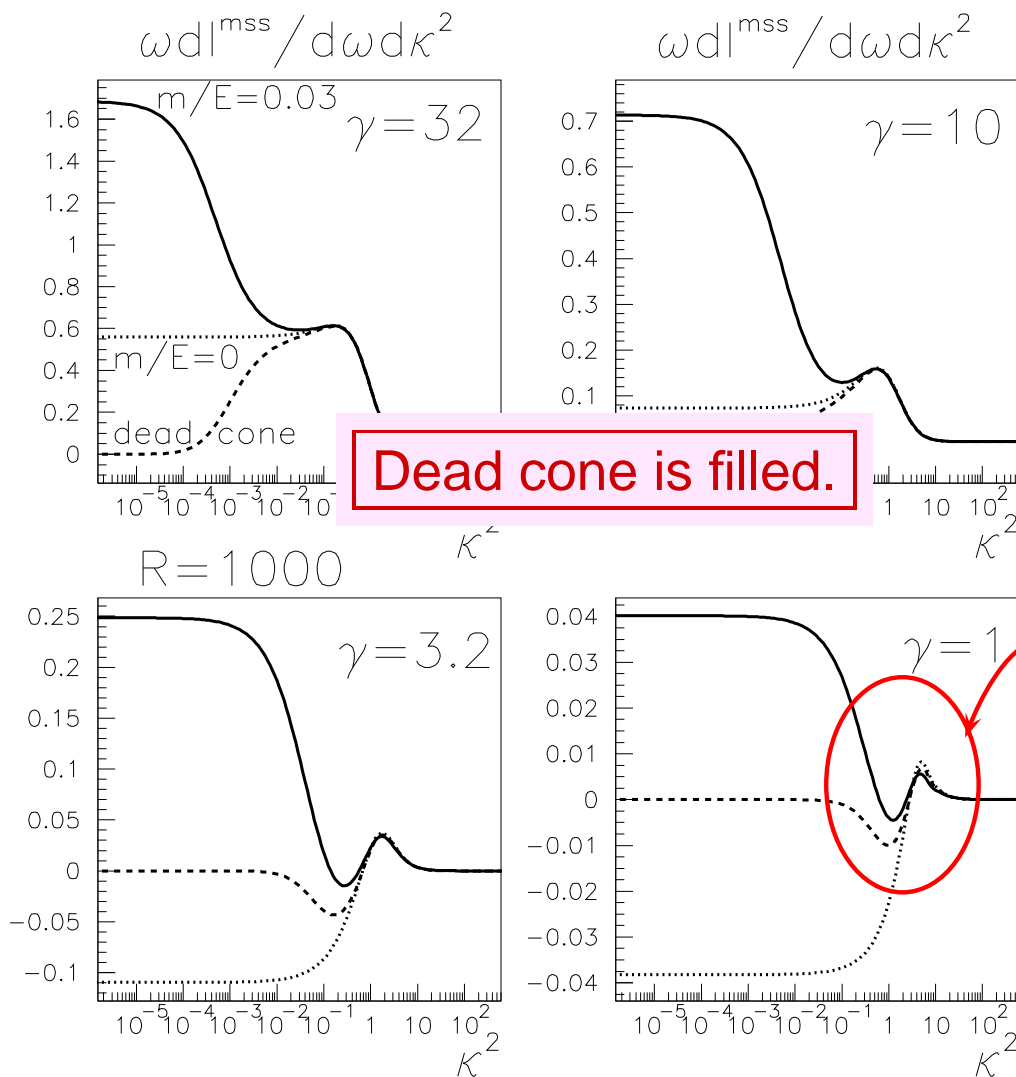


- $\kappa^2 = \frac{k_{\perp}^2}{\hat{q}L}$, $\omega_c = \hat{q}L^2/2$,
 $R = \omega_c L$, $\gamma = \omega_c/\omega$,
 $M^2 = \frac{x^2 m^2}{\hat{q}L}$.

- Dead cone \equiv massless
 $\times F(\kappa, M) = \left(\frac{\kappa^2}{\kappa^2 + M^2} \right)^2$.
- $I_{\text{med}} + I_{\text{vac}} \geq 0$.
- For $m > 0$, $I_{\text{med}} \geq 0$.

2. Results (I) – the dead cone is filled:

Néstor Armesto



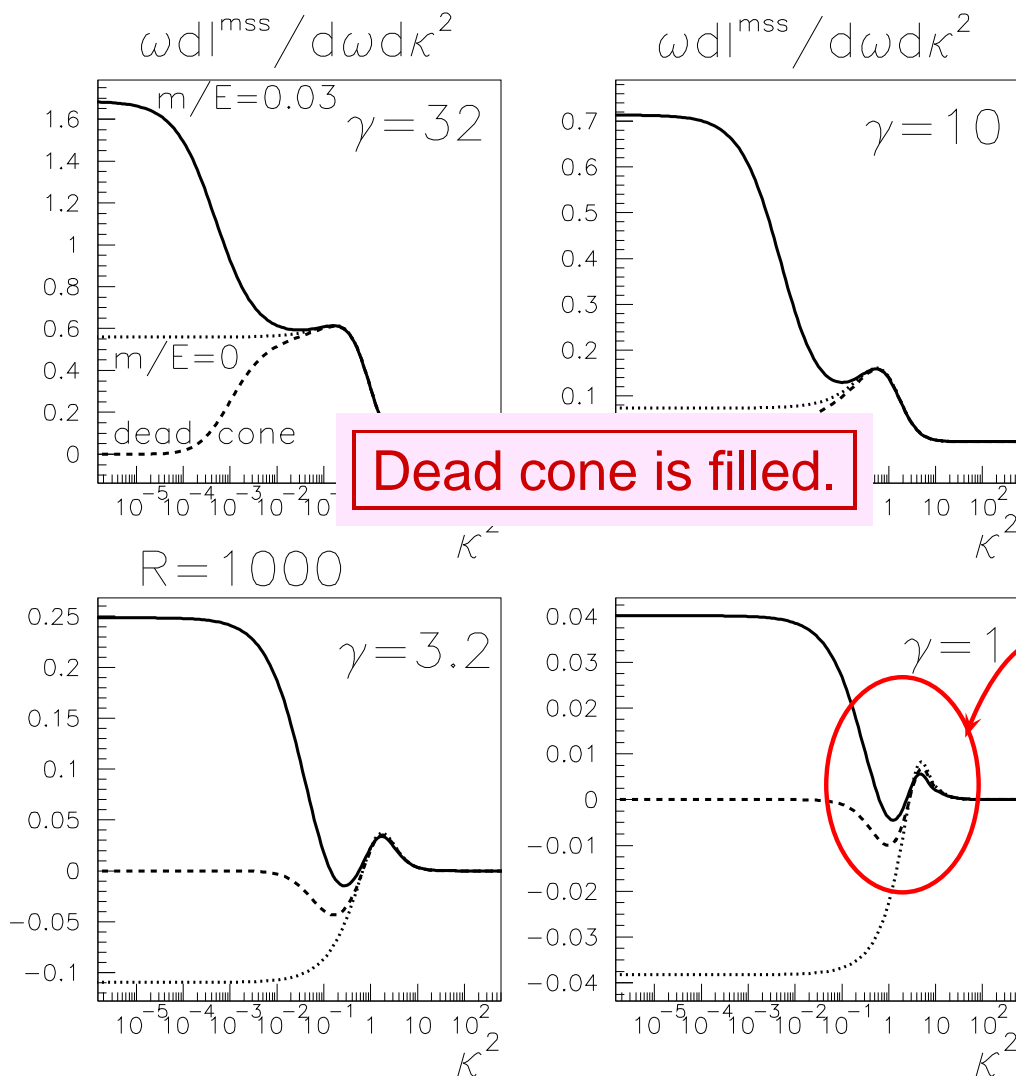
- $\kappa^2 = \frac{k_{\perp}^2}{\hat{q}L}$, $\omega_c = \hat{q}L^2/2$,
 $R = \omega_c L$, $\gamma = \omega_c/\omega$,
 $M^2 = \frac{x^2 m^2}{\hat{q}L}$.

- Dead cone \equiv massless
 $\times F(\kappa, M) = \left(\frac{\kappa^2}{\kappa^2 + M^2} \right)^2$.
- $I_{\text{med}} + I_{\text{vac}} \geq 0$.
- For $m > 0$, $I_{\text{med}} \geq 0$.

- At large κ , more radiation from massless than from massive quarks.

2. Results (I) – the dead cone is filled:

Néstor Armesto



- $\kappa^2 = \frac{k_{\perp}^2}{\hat{q}L}$, $\omega_c = \hat{q}L^2/2$,
 $R = \omega_c L$, $\gamma = \omega_c/\omega$,
 $M^2 = \frac{x^2 m^2}{\hat{q}L}$.

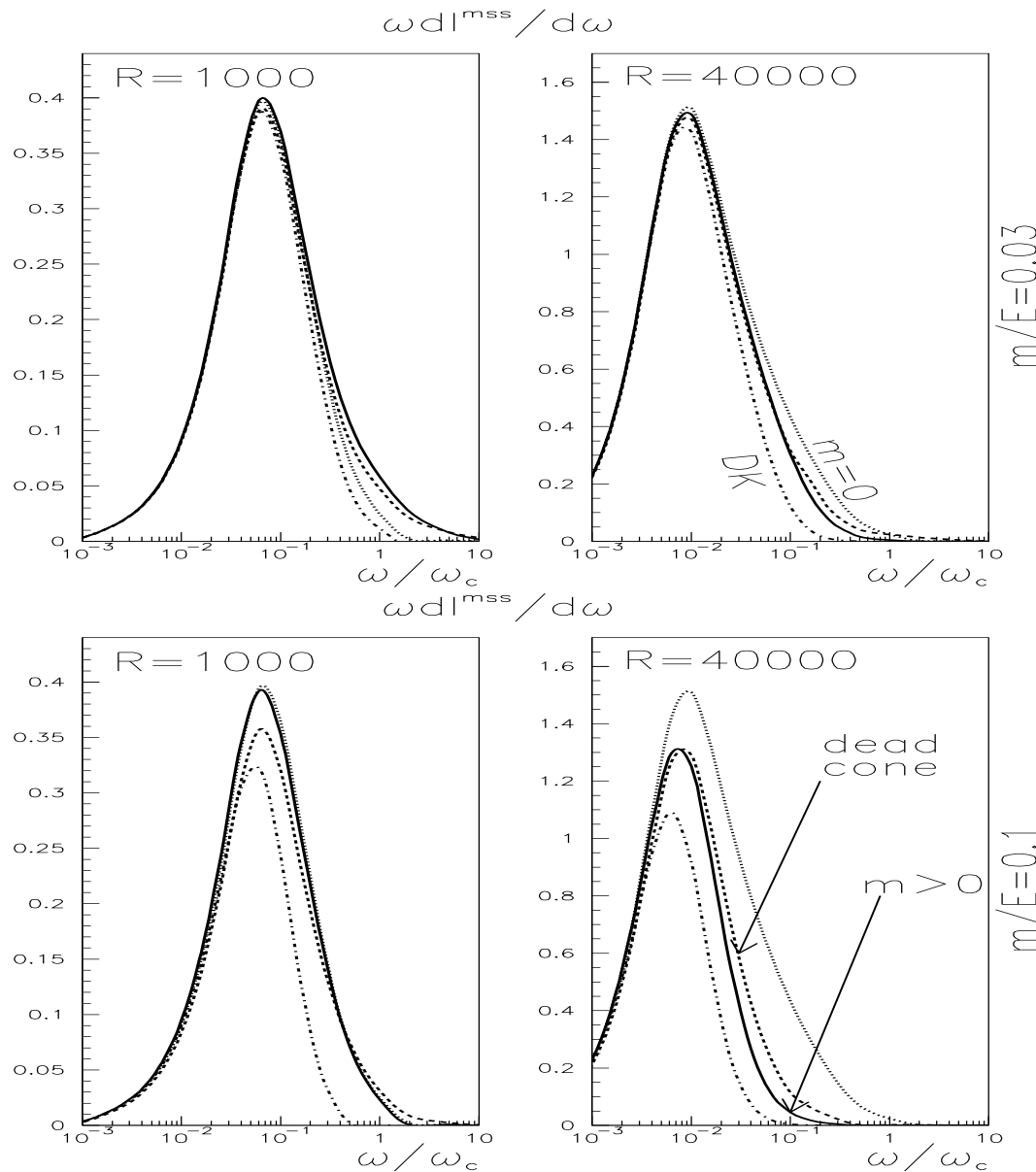
- Dead cone \equiv massless
 $\times F(\kappa, M) = \left(\frac{\kappa^2}{\kappa^2 + M^2} \right)^2$.
- $I_{\text{med}} + I_{\text{vac}} \geq 0$.
- For $m > 0$, $I_{\text{med}} \geq 0$.

- At large κ , more radiation from massless than from massive quarks.

- Kinematically, the dead cone is only a small region of the available phase space.

2. Results (II) – k_{\perp} -integrated radiation is suppressed:

Néstor Armesto

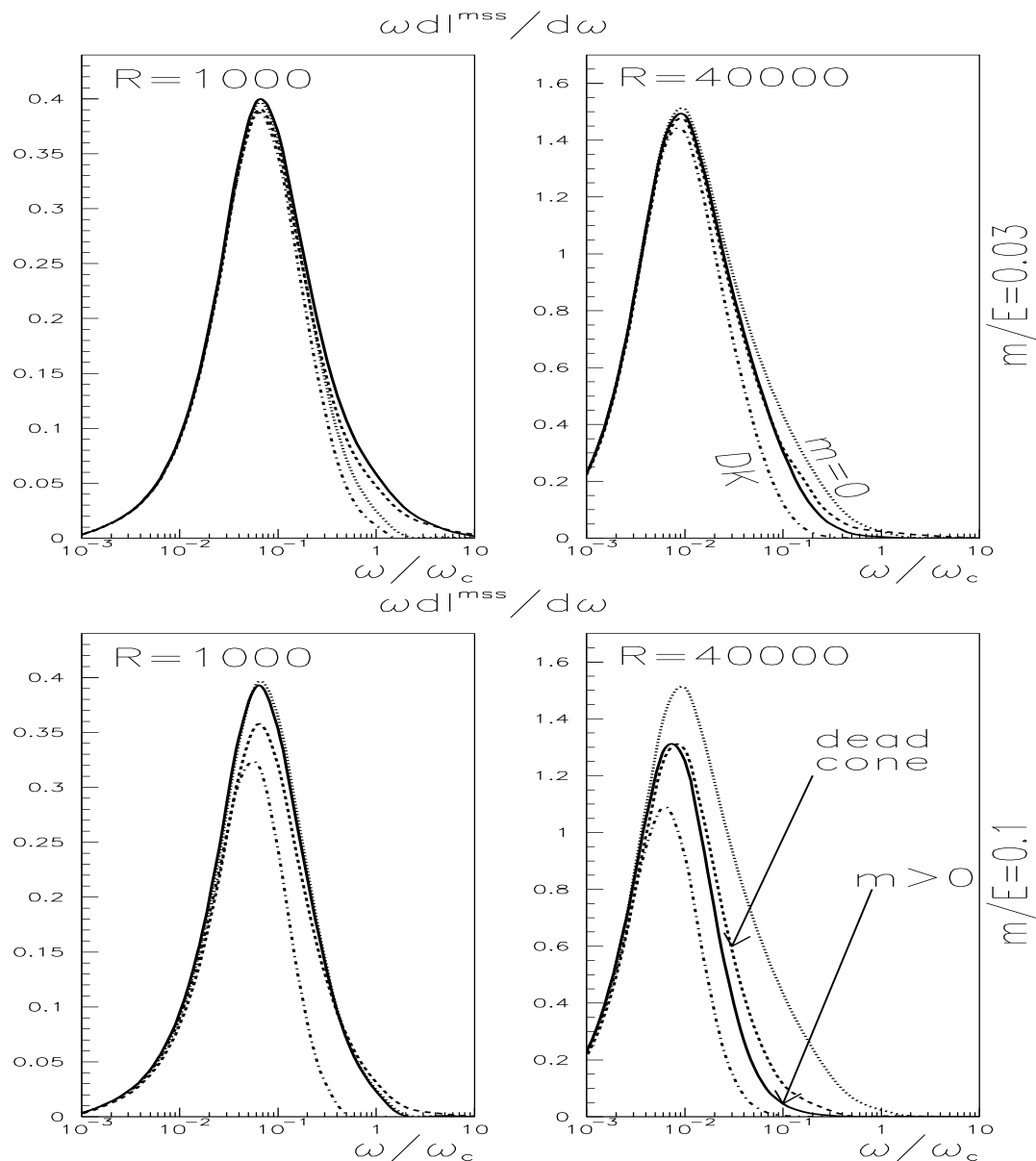


- $\kappa^2 = \frac{k_{\perp}^2}{\hat{q}L}$, $\omega_c = \hat{q}L^2/2$,
 $R = \omega_c L$, $\gamma = \omega_c/\omega$,
 $M^2 = \frac{x^2 m^2}{\hat{q}L}$.

- $\omega \frac{dI}{d\omega} = \int_0^{\omega} dk_{\perp} \omega \frac{dI}{d\omega dk_{\perp}}$.

2. Results (II) – k_{\perp} -integrated radiation is suppressed:

Néstor Armesto



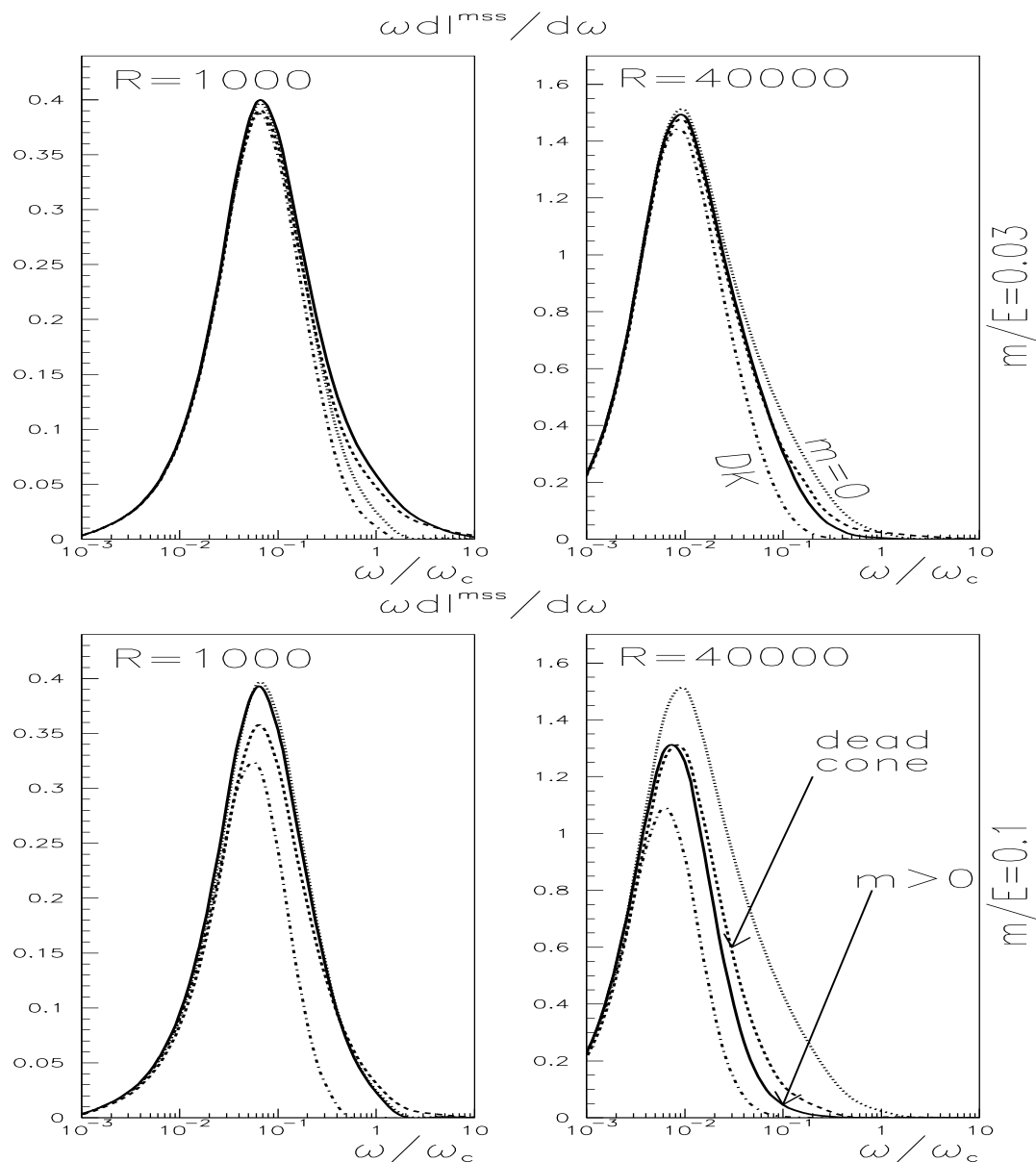
- $\kappa^2 = \frac{k_{\perp}^2}{\hat{q}L}$, $\omega_c = \hat{q}L^2/2$,
 $R = \omega_c L$, $\gamma = \omega_c/\omega$,
 $M^2 = \frac{x^2 m^2}{\hat{q}L}$.

- $\omega \frac{dI}{d\omega} = \int_0^{\omega} d\mathbf{k}_{\perp} \omega \frac{dI}{d\omega d\mathbf{k}_{\perp}}$.

- k_{\perp} -integrated radiation is suppressed (qualitatively in agreement with dead cone proposal).

2. Results (II) – k_{\perp} -integrated radiation is suppressed:

Néstor Armesto



- $\kappa^2 = \frac{k_{\perp}^2}{\hat{q}L}$, $\omega_c = \hat{q}L^2/2$,
 $R = \omega_c L$, $\gamma = \omega_c/\omega$,
 $M^2 = \frac{x^2 m^2}{\hat{q}L}$.

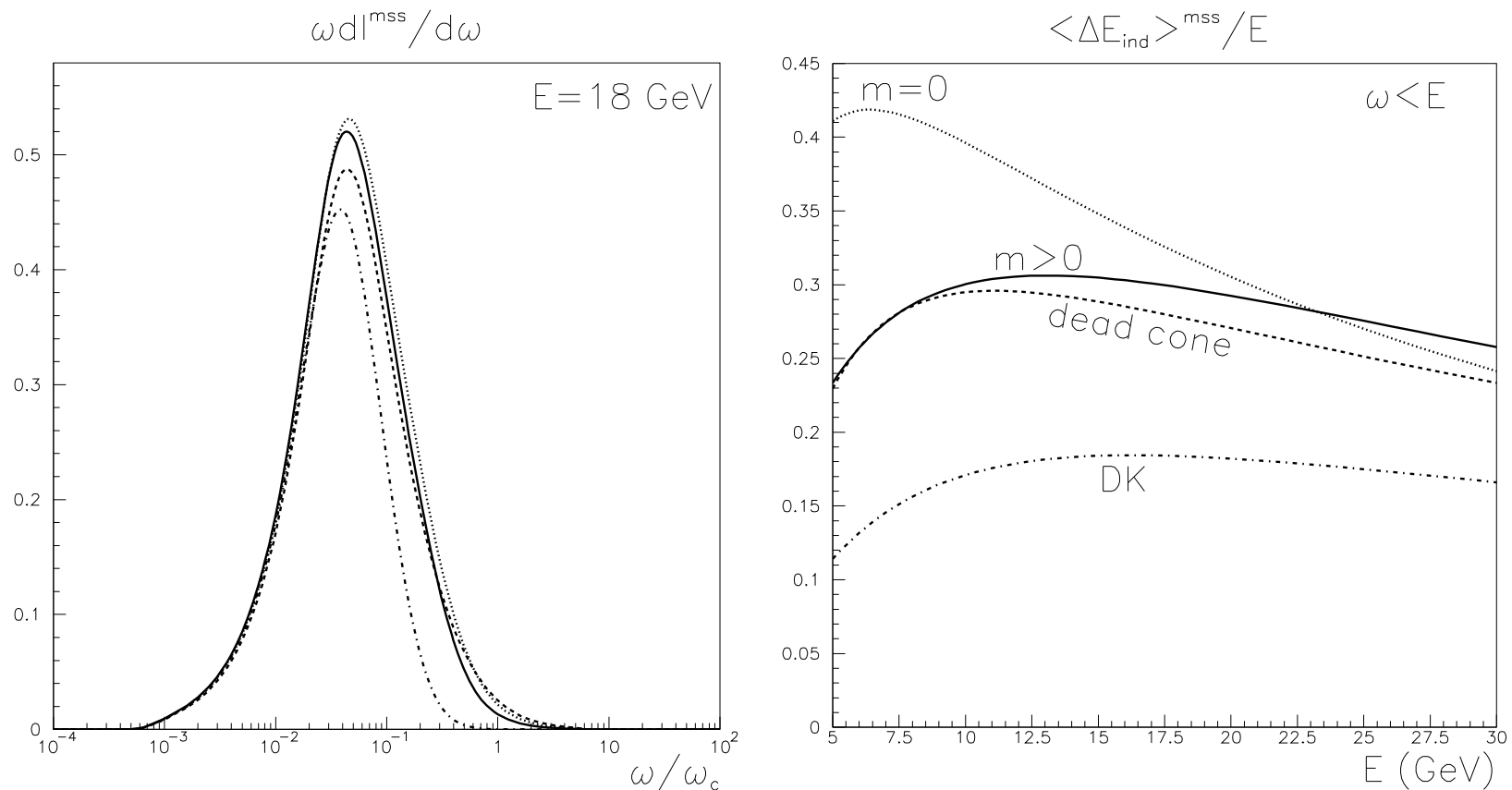
- $\omega \frac{dI}{d\omega} = \int_0^{\omega} d\mathbf{k}_{\perp} \omega \frac{dI}{d\omega d\mathbf{k}_{\perp}}$.

- k_{\perp} -integrated radiation is suppressed (qualitatively in agreement with dead cone proposal).

- $F(\kappa, M) = \left(\frac{\kappa^2}{\kappa^2 + M^2} \right)^2$.

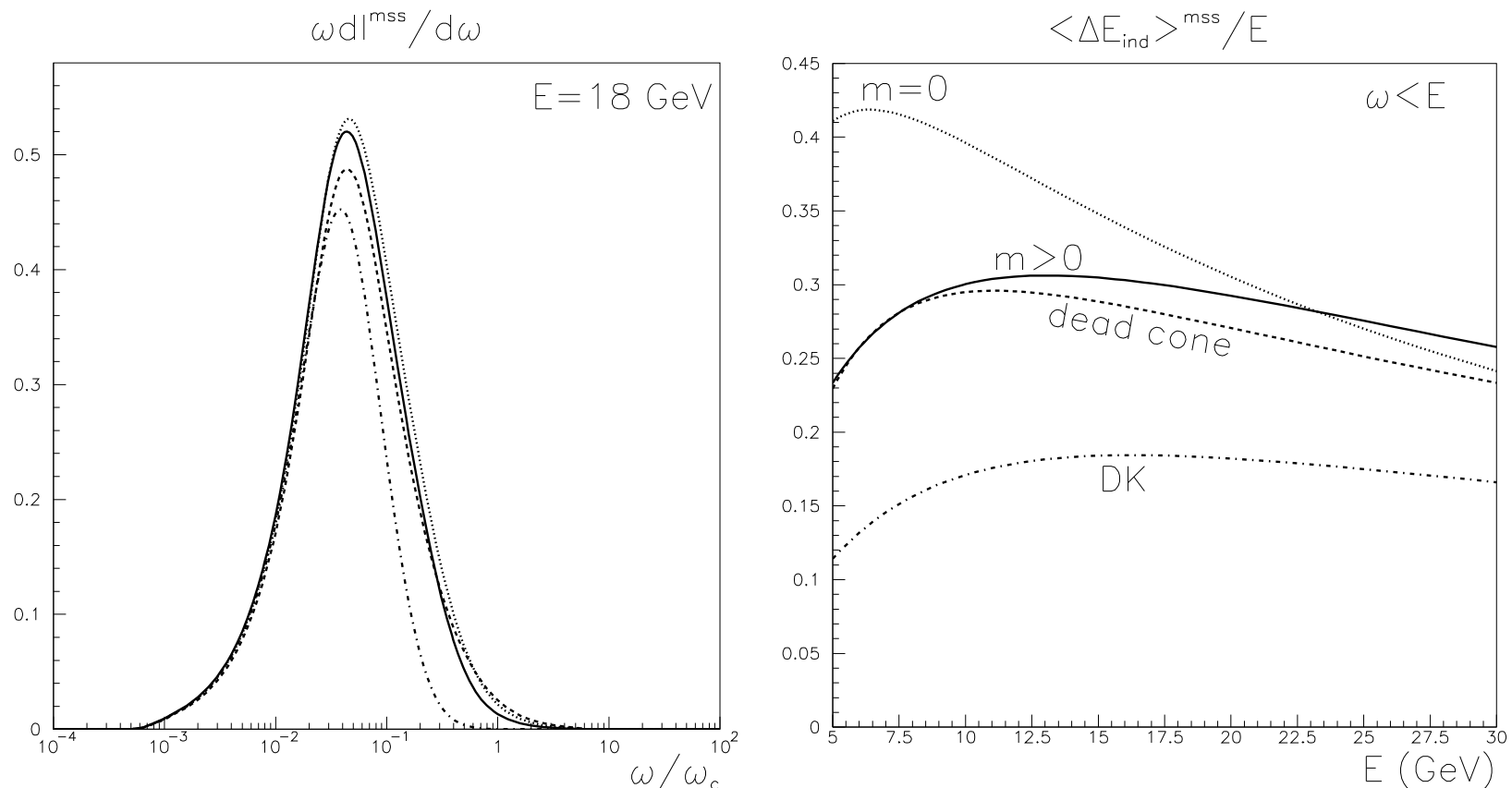
- $F_{DK}(M) = F(\mathbf{k}_{\perp}^2 = \sqrt{\hat{q}\omega}, M)$ (Dokshitzer, Kharzeev, PLB519(01)199) underestimates quantitatively the emission.

2. Results (III) – uncertainties in application to RHIC and LHC: Néstor Armesto



- $\langle \Delta E_{\text{ind}} \rangle = \int_0^E d\omega \omega \frac{dI_{\text{medium}}}{d\omega}$; $L = 6$ fm, $\hat{q} = 0.8$ GeV²/fm
 ($R = \omega_c L = 2000$, $\omega_c = \hat{q} L^2 / 2 = 67.5$ GeV) (Salgado, Wiedemann, PRD68(03)014008) and $m = 1.5$ GeV.

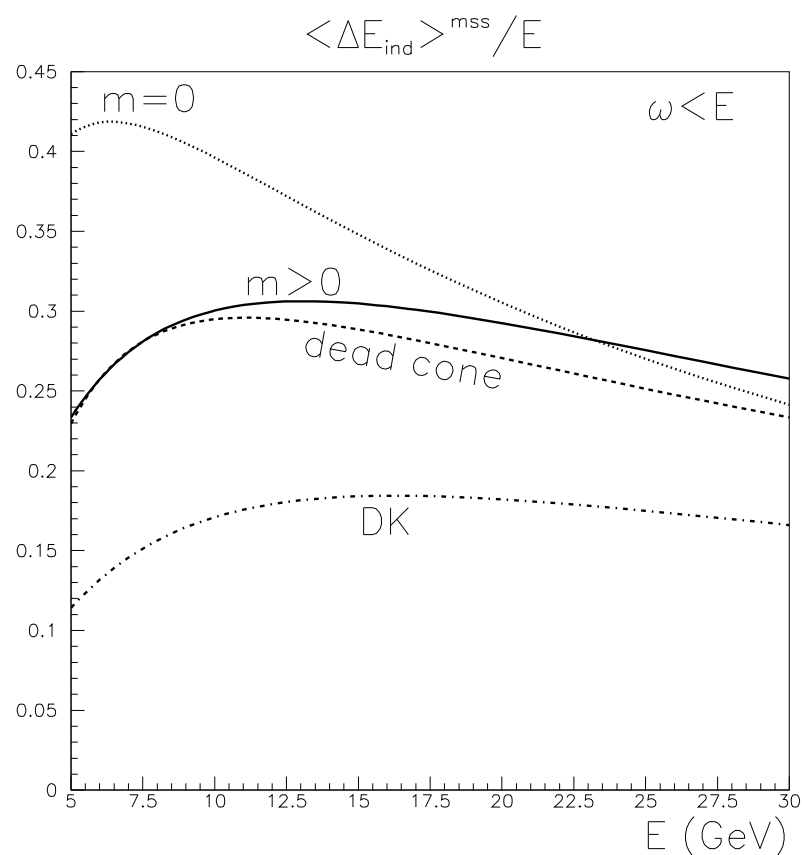
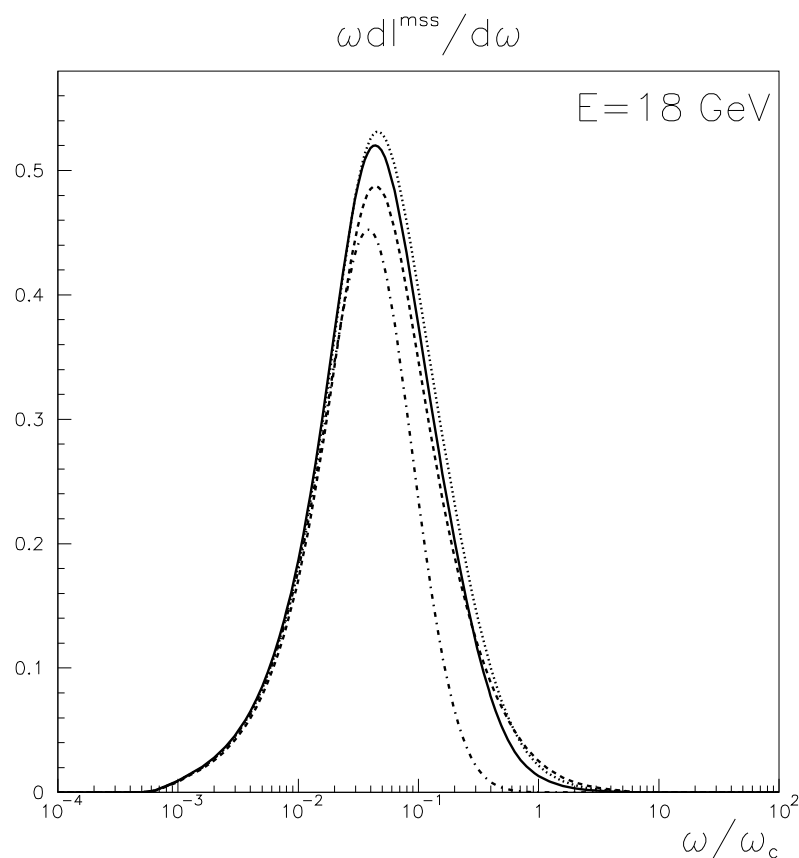
2. Results (III) – uncertainties in application to RHIC and LHC: Néstor Armesto



- $\langle \Delta E_{\text{ind}} \rangle = \int_0^E d\omega \omega \frac{dI_{\text{medium}}}{d\omega}$; $L = 6 \text{ fm}$, $\hat{q} = 0.8 \text{ GeV}^2/\text{fm}$
 $(R = \omega_c L = 2000, \omega_c = \hat{q} L^2 / 2 = 67.5 \text{ GeV})$ (Salgado, Wiedemann, PRD68(03)014008) and $m = 1.5 \text{ GeV}$.

- Energy loss for charm at RHIC turns out to be a factor ~ 2 smaller than for light quarks, but should still be observable.

2. Results (III) – uncertainties in application to RHIC and LHC: Néstor Armesto



- $\langle \Delta E_{\text{ind}} \rangle = \int_0^E d\omega \omega \frac{dI_{\text{medium}}}{d\omega}$; $L = 6 \text{ fm}$, $\hat{q} = 0.8 \text{ GeV}^2/\text{fm}$
 $(R = \omega_c L = 2000, \omega_c = \hat{q} L^2 / 2 = 67.5 \text{ GeV})$ (Salgado, Wiedemann, PRD68(03)014008) and $m = 1.5 \text{ GeV}$.

- Energy loss for charm at RHIC turns out to be a factor ~ 2 smaller than for light quarks, but should still be observable.
- **Caveat:** uncertainties in all computations are significant ...

3. Conclusions:

Néstor Armesto

♣ Medium-induced gluon radiation off massless and massive quarks is treated in the same formalism. Ours is the first k_{\perp} -differential result (consistent with available k_{\perp} -integrated ones (Djordjevic, Gyulassy, PLB560(03)37; NPA733(04)265; Zhang, Wang, Wang, nucl-th/0309040)).

- ♣ Medium-induced gluon radiation off massless and massive quarks is treated in the same formalism. Ours is the first k_{\perp} -differential result (consistent with available k_{\perp} -integrated ones (Djordjevic, Gyulassy, PLB560(03)37; NPA733(04)265; Zhang, Wang, Wang, nucl-th/0309040)).
- Medium-induced gluon radiation fills the dead cone.

- ♣ Medium-induced gluon radiation off massless and massive quarks is treated in the same formalism. Ours is the first k_{\perp} -differential result (consistent with available k_{\perp} -integrated ones (Djordjevic, Gyulassy, PLB560(03)37; NPA733(04)265; Zhang, Wang, Wang, nucl-th/0309040)).
- Medium-induced gluon radiation fills the dead cone.
- But the dead cone (i.e. the low- k_{\perp} region) does not dominate the energy loss.

♣ Medium-induced gluon radiation off massless and massive quarks is treated in the same formalism. Ours is the first k_{\perp} -differential result (consistent with available k_{\perp} -integrated ones (Djordjevic, Gyulassy, PLB560(03)37; NPA733(04)265; Zhang, Wang, Wang, nucl-th/0309040)).

- Medium-induced gluon radiation fills the dead cone.

- But the dead cone (i.e. the low- k_{\perp} region) does not dominate the energy loss.

♣ Our study suggests that energy loss for charmed hadrons at RHIC should be smaller than that for lighter hadrons, but still sizable (for $p_{\perp}^{\text{hadron}} \simeq 5 \div 10 \text{ GeV}/c$).

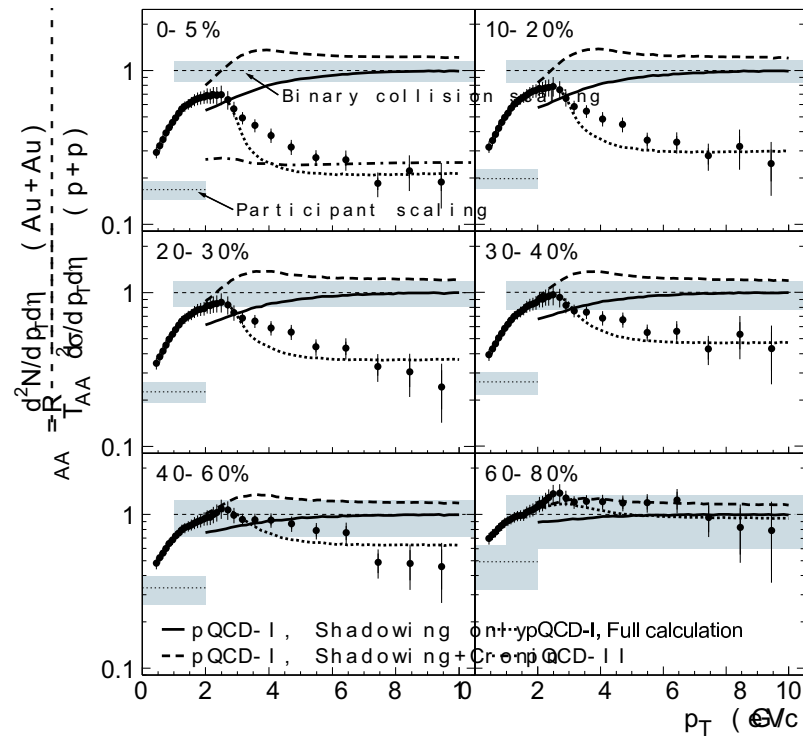
♣ Medium-induced gluon radiation off massless and massive quarks is treated in the same formalism. Ours is the first k_{\perp} -differential result (consistent with available k_{\perp} -integrated ones (Djordjevic, Gyulassy, PLB560(03)37; NPA733(04)265; Zhang, Wang, Wang, nucl-th/0309040)).

- Medium-induced gluon radiation fills the dead cone.

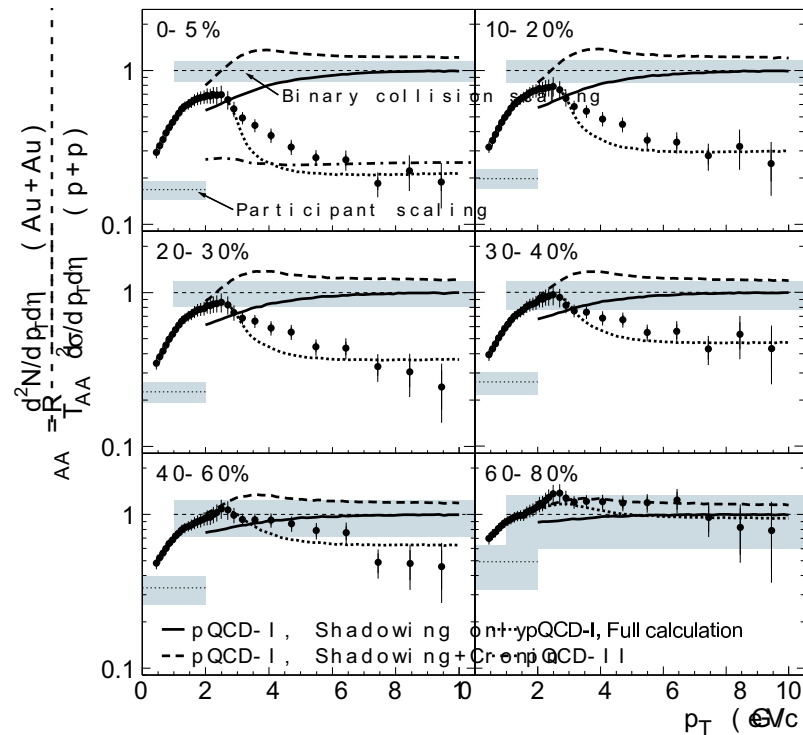
- But the dead cone (i.e. the low- k_{\perp} region) does not dominate the energy loss.

♣ Our study suggests that energy loss for charmed hadrons at RHIC should be smaller than that for lighter hadrons, but still sizable (for $p_{\perp}^{\text{hadron}} \simeq 5 \div 10 \text{ GeV}/c$).

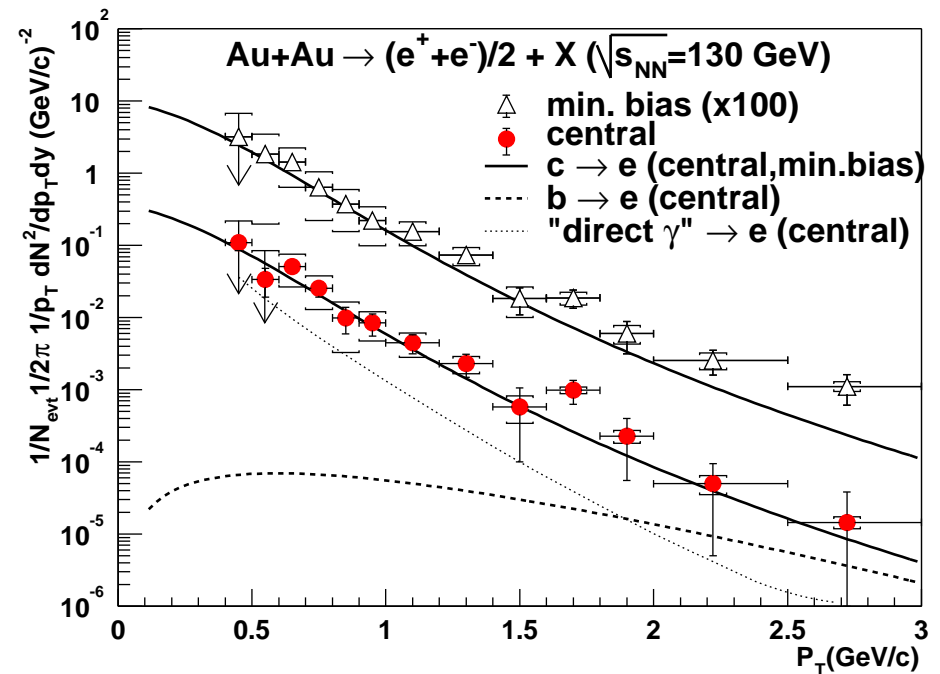
- Uncertainties motivate the computation of $1/E$ corrections.



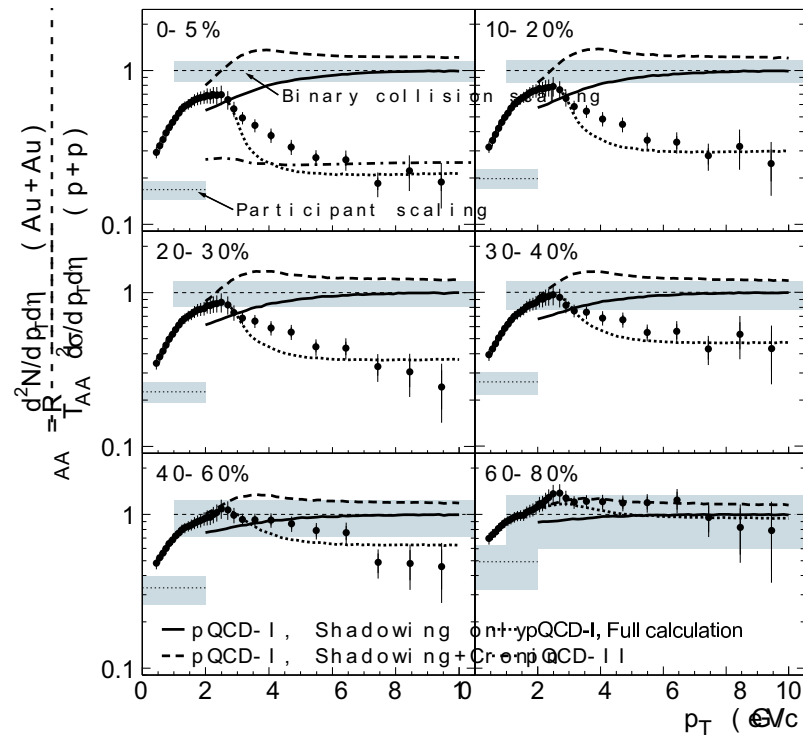
STAR: PRL91(03)172302, AuAu@200 GeV, charged.



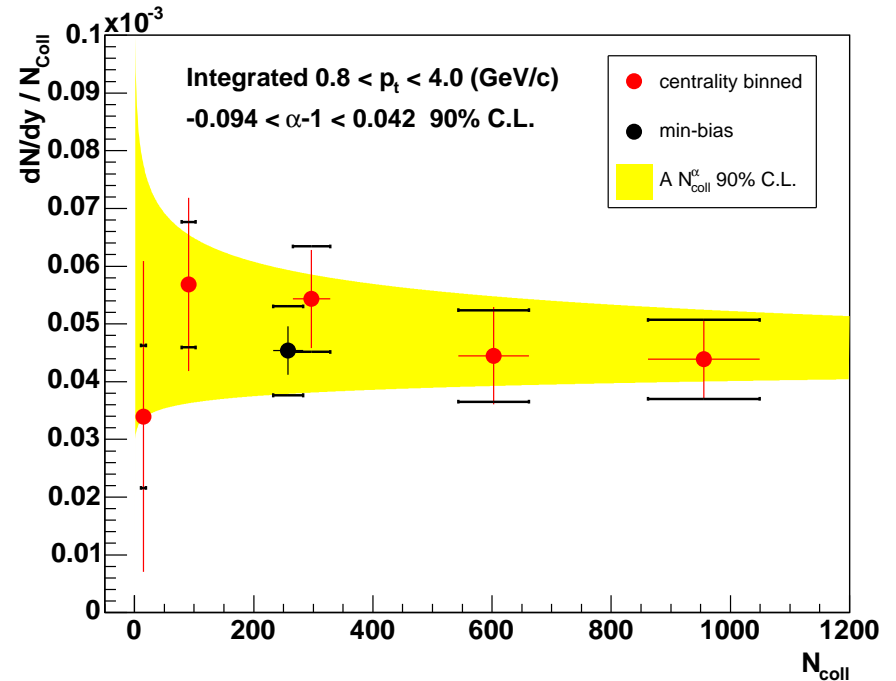
STAR: PRL91(03)172302, AuAu@200 GeV, charged.



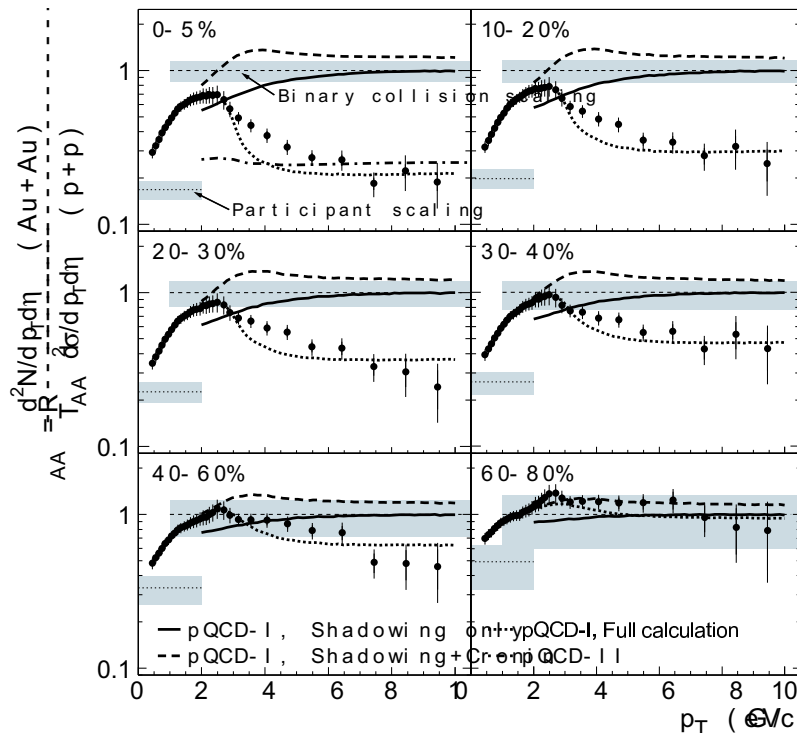
PHENIX: PRL88(02)192303.



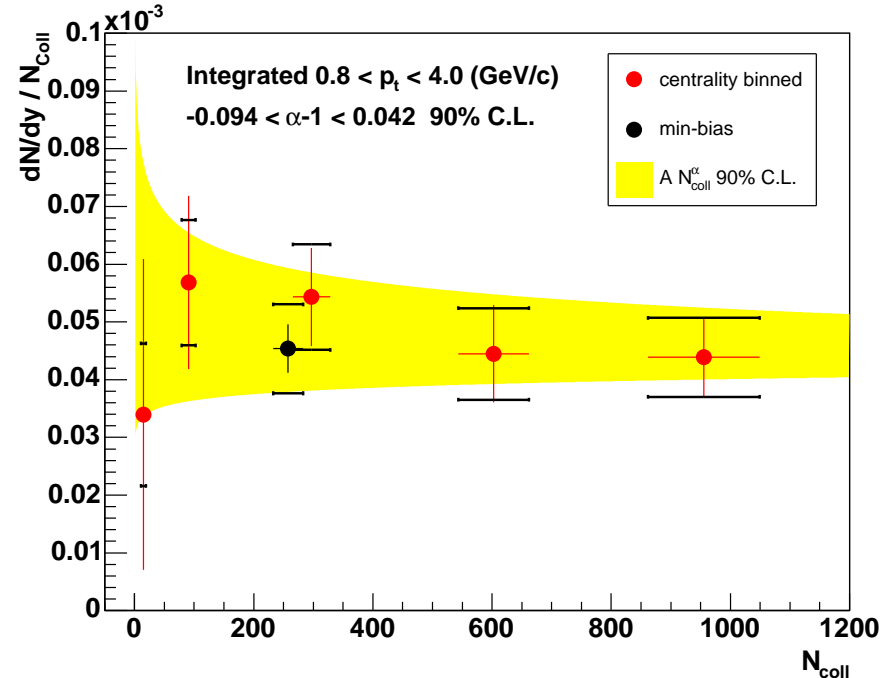
STAR: PRL91(03)172302, AuAu@200 GeV, charged.



PHENIX: Kelly, nucl-ex/0403057.



STAR: PRL91(03)172302, AuAu@200 GeV, charged.



PHENIX: Kelly, nucl-ex/0403057.

⇒ apparently less suppression, if any, for charmed than for lighter hadrons (but single e^- spectra do not constrain energy loss significantly (Batsouli, Kelly, Gyulassy, Nagle, PLB557(03)26), and p_{\perp} are small so hadronization effects have to be taken into account).