

# Measurements of low energy $e^+e^-$ hadronic cross sections and implications for the muon $g-2$

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( Representing the BABAR Collaboration )



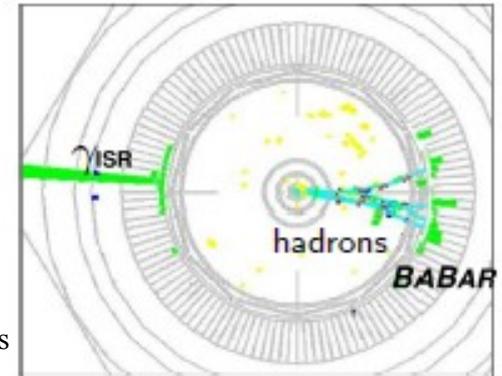
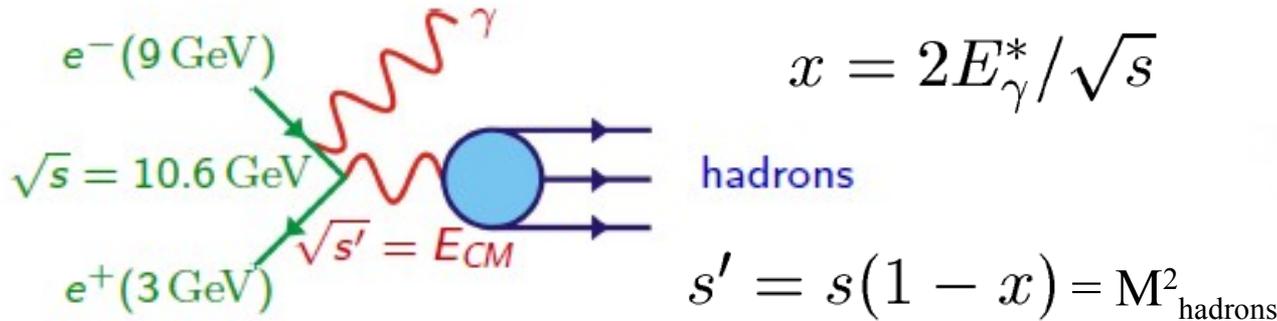
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# Outline

- The BABAR ISR (Initial State Radiation)  $\pi\pi$  analysis
  - Test of the method:  $e^+e^- \rightarrow \mu^+\mu^-(\gamma)$
  - Results on  $e^+e^- \rightarrow \pi^+\pi^-(\gamma)$
  - Results on  $K^+K^-(\gamma)$
  - Conclusions
- PRL 103, 231801 (2009)  
PRD 86, 032013 (2012)  
PRD 88, 032013 (2013)

# The ISR method at BABAR



- High energy ( $E_{\gamma}^* > 3 \text{ GeV}$ ) detected at large angle  
 → defines  $\sqrt{s'} = E_{CM}$  and provides strong background rejection
- Event topology: ISR photon back-to-back to hadrons  
 → high acceptance, large boost to hadrons (measurements from threshold and easier PID)
- Final state can be hadronic or leptonic (QED)  
 →  $\mu^+\mu^-\gamma(\gamma)$  events used to get ISR luminosity
- Kinematic fit including ISR photon  
 → removes multihadronic background; improves mass resolution (a few MeV)
- Continuous measurement from threshold to 3-5 GeV  
 → reduced systematic uncertainties compared to multiple data sets with different colliders and detectors

# The BABAR ISR program

- cover an almost complete set of significant exclusive  $e^+e^-$  annihilation channels up to 2 GeV

- published:

$K^+ K^-$  PRD 2013

$\pi^+\pi^-$  PRL 2009; PRD 2012

$\pi^+\pi^-\pi^0$  PRD 2004

$\Phi f^0(980)$  PRD 2006; PRD 2007

$p \bar{p}$  PRD 2006; PRD 2013

$\Lambda \Lambda, \Lambda \Sigma^0, \Sigma^0 \Sigma^0$  PRD 2007

$2(\pi^+\pi^-), K^+K^-\pi^+\pi^-, K^+K^-2\pi^0, 2(K^+K^-)$  PRD 2007; PRD 2012

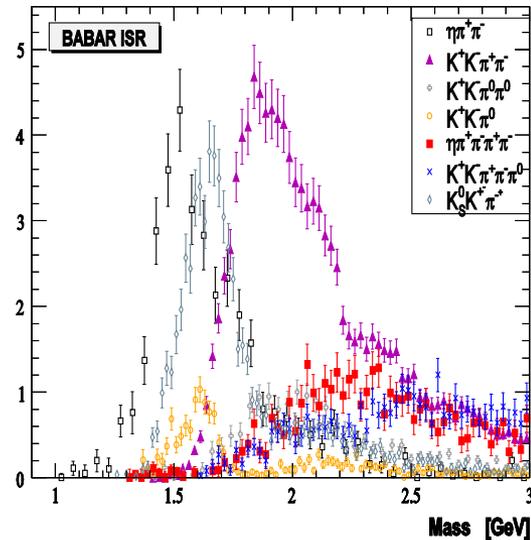
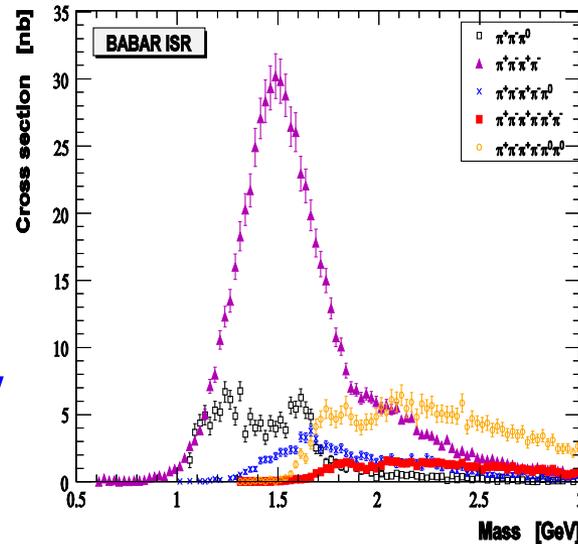
$K_S^0 K^+\pi^-, K^+K^-\pi^0, K^+K^-\eta$  PRD 2005; PRD 2008

$2(\pi^+\pi^-)\pi^0, 2(\pi^+\pi^-)\eta, K^+K^-\pi^+\pi^-\pi^0, K^+K^-\pi^+\pi^-\eta$  PRD 2007

$3(\pi^+\pi^-), 2(\pi^+\pi^-\pi^0), 2(\pi^+\pi^-)K^+K^-$  PRD 2006

- in progress:

$\pi^+\pi^-2\pi^0, K_S^0 K_L^0, K_S^0 K_L^0 \pi^+\pi^-, K_S^0 K^+\pi^+\pi^0, K_S^0 K^+\pi^+\eta$

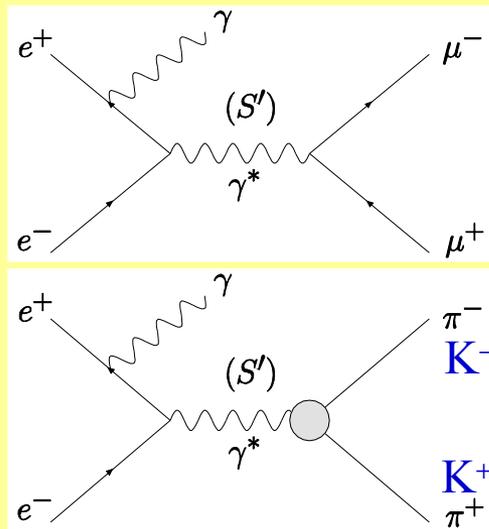


# The relevant processes for the $\pi\pi$ and KK measurements

$e^+e^- \rightarrow \mu^+\mu^- \gamma_{\text{ISR}} (\gamma_{\text{add.}}), \pi^+\pi^- \gamma_{\text{ISR}} (\gamma_{\text{add.}})$  and  $K^+K^- \gamma_{\text{ISR}} (\gamma_{\text{add.}})$

measured simultaneously

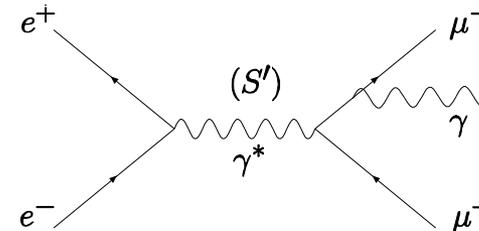
ISR



$$x = 2E_{\gamma}^*/\sqrt{s}$$

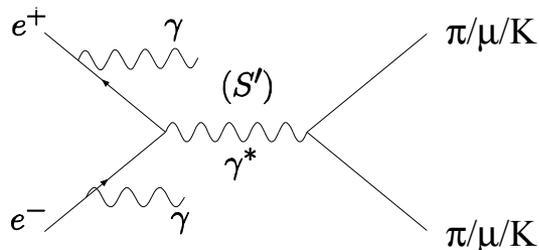
$$s' = s(1 - x)$$

FSR

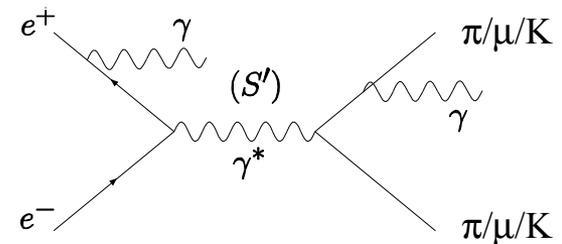


LO FSR negligible for  $\pi\pi$  and KK at  $s \sim (10.6 \text{ GeV})^2$

ISR + additional ISR

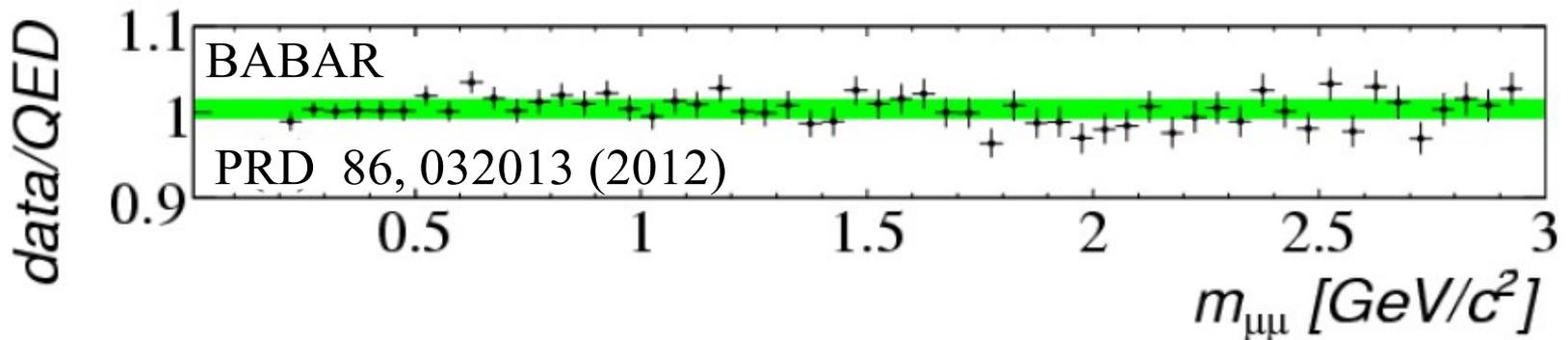


ISR + additional FSR



# QED test with $\mu\mu\gamma$ sample

- absolute comparison of  $\mu\mu$  mass spectra in data and in simulation
- simulation corrected for data/MC efficiencies
- AfkQed corrected for incomplete NLO using Phokhara
- strong test (ISR probability function drops out for  $\pi\pi$  cross section)



$$\frac{\sigma_{\mu\mu\gamma(\gamma)}^{data}}{\sigma_{\mu\mu\gamma(\gamma)}^{NLO\ QED}} = 1 + (4.0 \pm 1.9 \pm 5.5 \pm 9.4) \times 10^{-3} \quad (0.2 - 3\ \text{GeV})$$

ISR  $\gamma$  efficiency 3.4 syst.  
trig/track/PID 4.0

BABAR  $e^+e^-$  luminosity

# Obtaining the $\pi\pi(\gamma_{\text{FSR}})$ cross section

$$\frac{dN_{\pi\pi\gamma(\gamma)}}{d\sqrt{s'}} = \frac{dL_{ISR}^{eff}}{d\sqrt{s'}} \epsilon_{\pi\pi\gamma(\gamma)}(\sqrt{s'}) \sigma_{\pi\pi(\gamma)}^0(\sqrt{s'})$$

Unfolded spectrum

Acceptance from MC + data/MC corrections

Effective ISR luminosity from  $\mu\mu\gamma(\gamma)$  analysis (similar equation + QED)

$\pi\pi$  mass spectrum unfolded (B. M. arXiv:0907.3791) for detector response

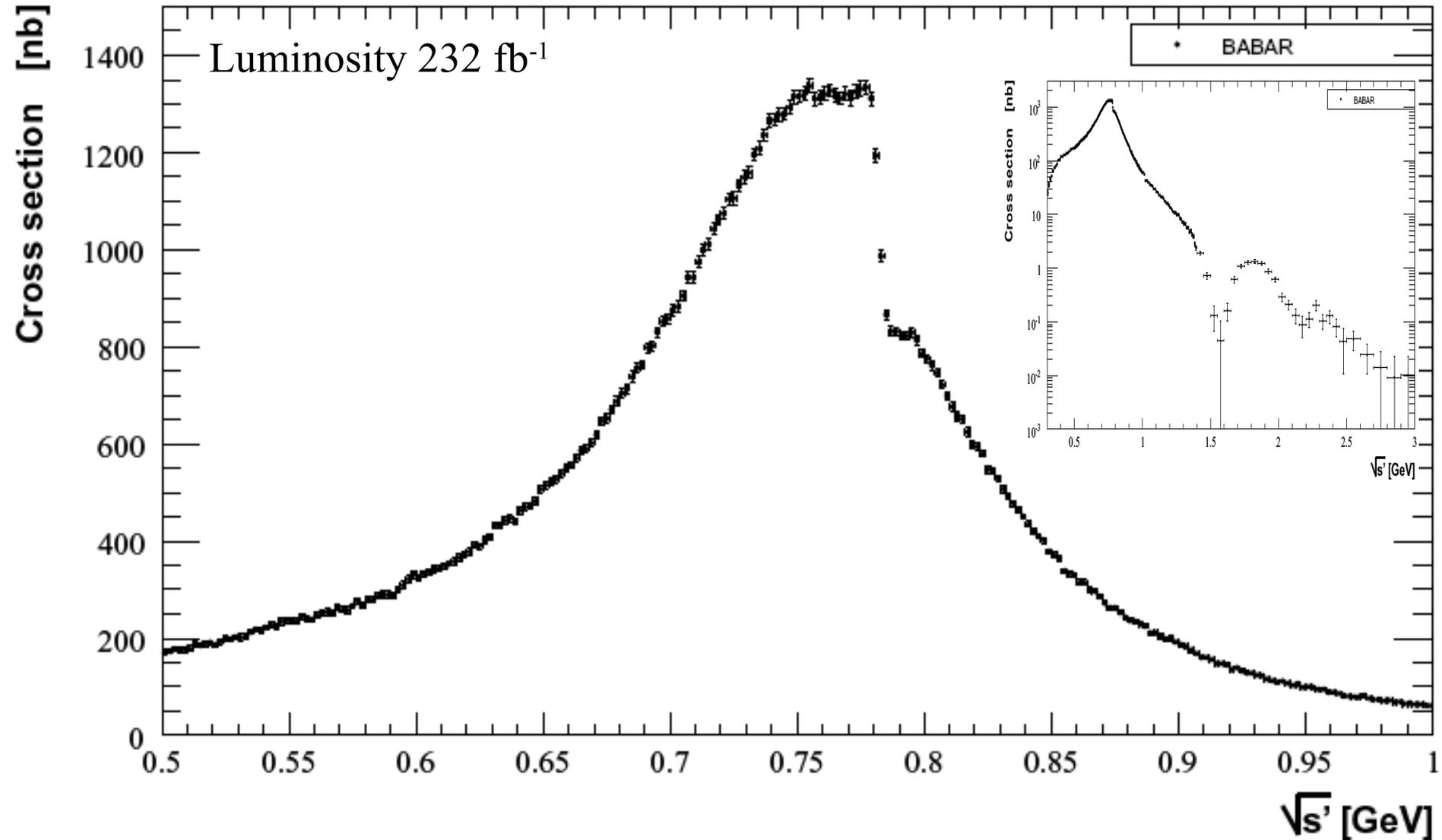
Additional ISR almost cancels in the procedure ( $\pi\pi\gamma(\gamma) / \mu\mu\gamma(\gamma)$  spectra ratio)  
Correction  $(2.5 \pm 1.0) \times 10^{-3} \Rightarrow$   **$\pi\pi$  cross section does not rely on accurate description of NLO in the MC generator**

ISR luminosity from  $\mu\mu\gamma(\gamma)$  in 50-MeV energy intervals  
(small compared to variation of efficiency corrections)

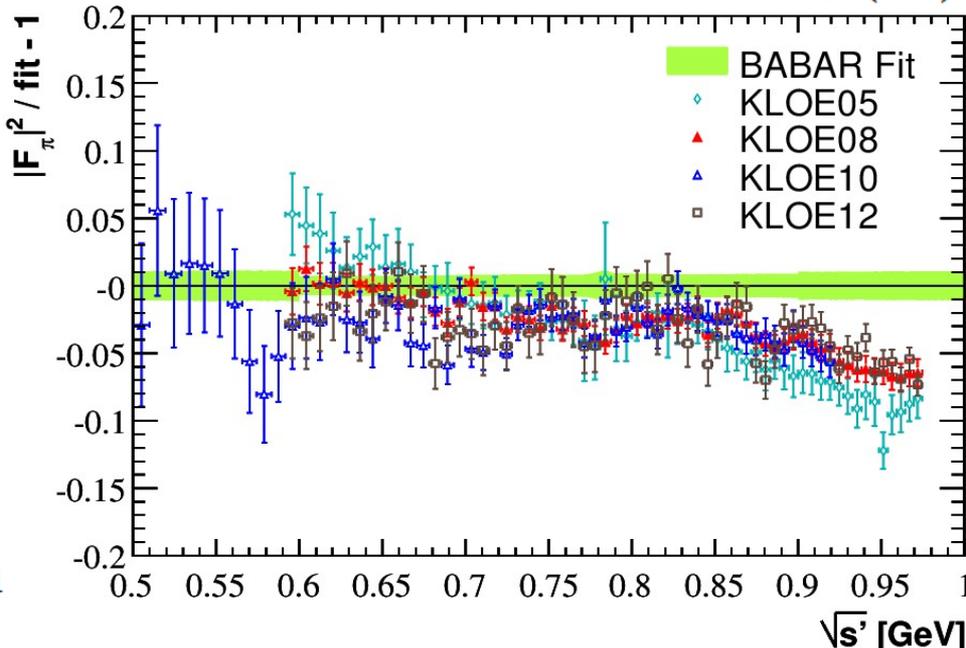
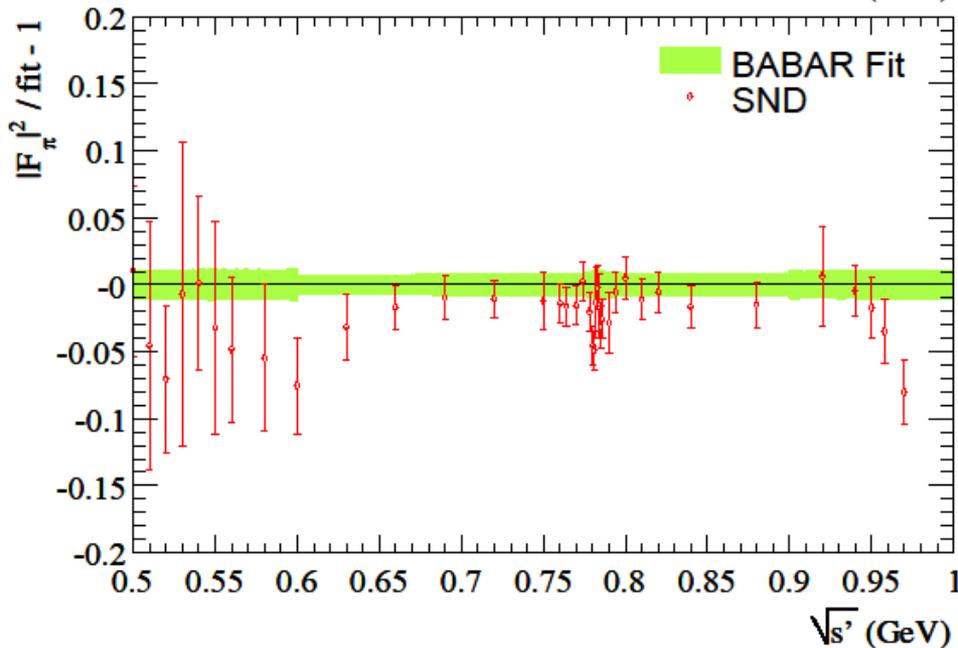
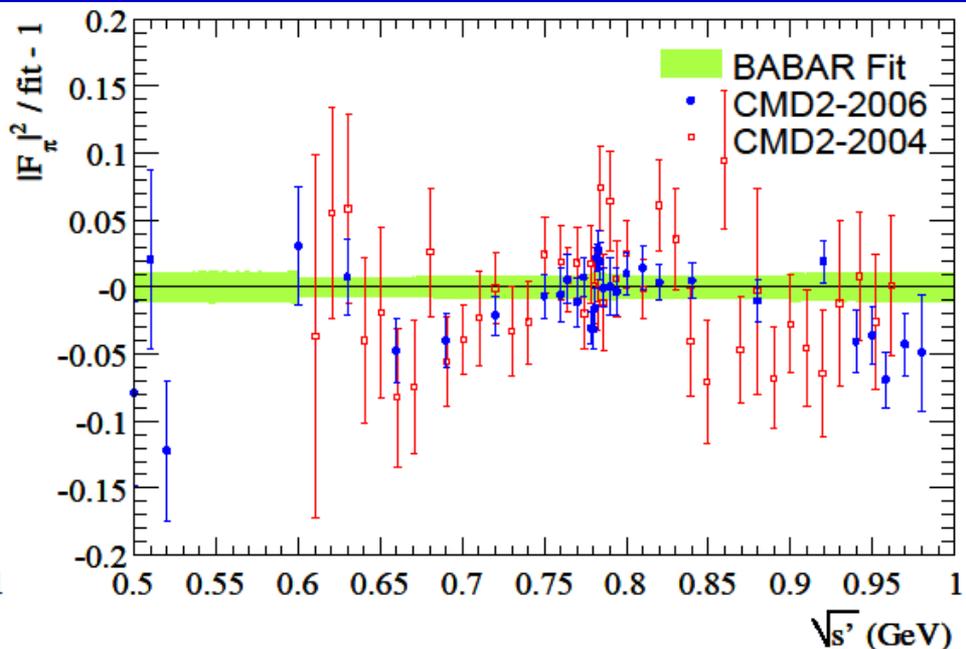
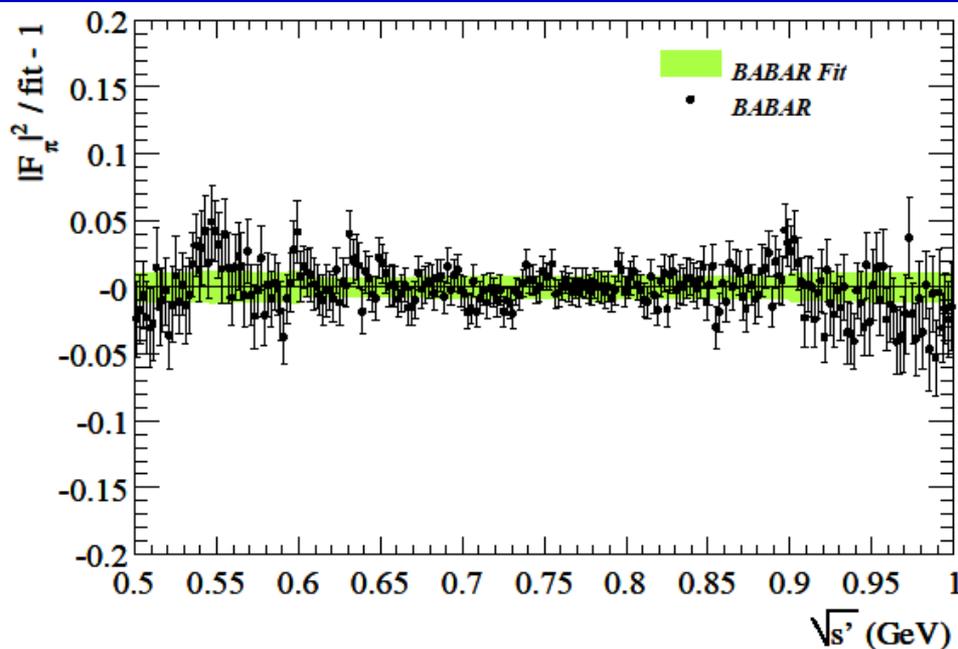
**Total systematic uncertainty of 0.5% in the  $\rho$  region (dominant for  $a_\mu$ )**

# BABAR results (PRL 103, 231801 (2009); PRD 86, 032013 (2012))

$e^+ e^- \rightarrow \pi^+ \pi^- (\gamma_{\text{FSR}})$  bare (no VP) cross section stat+syst uncertainties shown



# BABAR fit vs. $e^+e^-$ data (stat + syst errors included)



# Computing $a_\mu^{\pi\pi}$

$$a_\mu^{\pi\pi(\gamma),LO} = \frac{1}{4\pi^3} \int_{4m_\pi^2}^{\infty} ds K(s) \sigma_{\pi\pi(\gamma)}^0(s),$$

where  $K(s)$  is the QED kernel,

$$K(s) = x^2 \left(1 - \frac{x^2}{2}\right) + (1+x)^2 \left(1 + \frac{1}{x^2}\right) \left[ \ln(1+x) - x + \frac{x^2}{2} \right] + x^2 \frac{1+x}{1-x} \ln x$$

with  $x = (1 - \beta_\mu)/(1 + \beta_\mu)$  and  $\beta_\mu = (1 - 4m_\mu^2/s)^{1/2}$ .

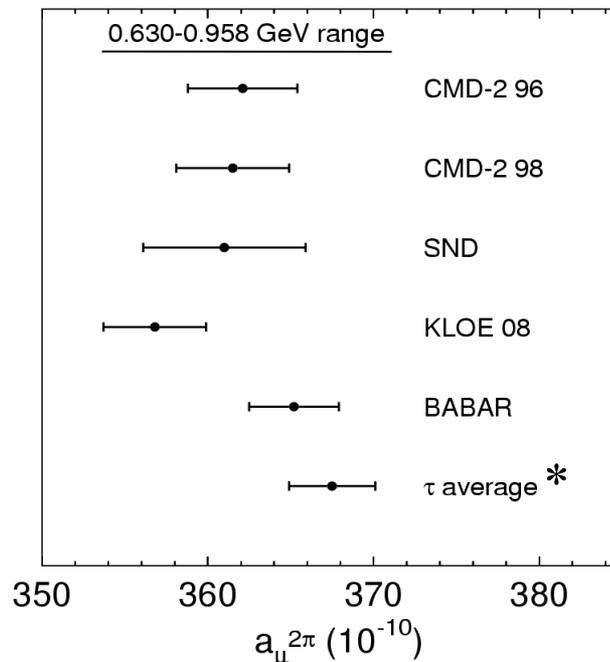
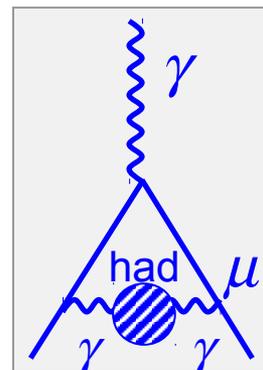
0.28–1.8 (GeV)

<b>BABAR</b>	$(514.1 \pm 3.8) \times 10^{-10}$
previous $e^+e^-$ combined	$(503.5 \pm 3.5) \times 10^{-10} *$
+ BABAR	$(508.4 \pm 2.9) \times 10^{-10} *$
+ KLOE 10	$(507.8 \pm 2.7) \times 10^{-10} *$
<b><math>\tau</math> combined</b>	$(515.2 \pm 3.5) \times 10^{-10} *$

Deviation between BNL measurement and theory prediction reduced using BABAR  $\pi^+\pi^-$  data

$$a_\mu [\text{exp}] - a_\mu [\text{SM}] = (19.8 \pm 8.4) \times 10^{-10} (2.4\sigma)$$

$\pi^+\pi^-$  from BABAR only



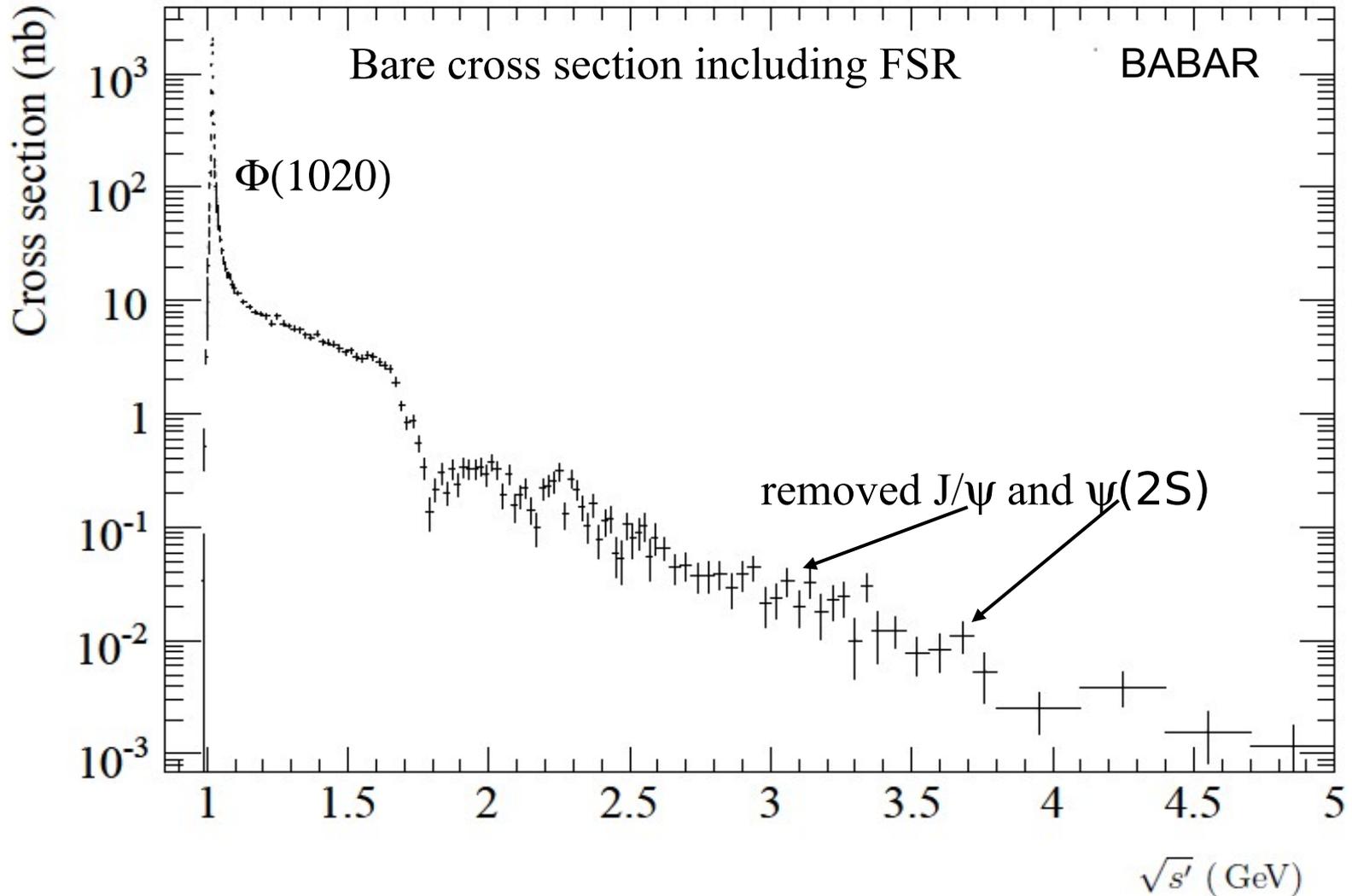
\* arXiv:0906.5443; 0908.4300; 1010.4180 M. Davier et al.

# Analysis of $e^+e^- \rightarrow K^+K^-(\gamma)$ (PRD 88, 032013 (2013))

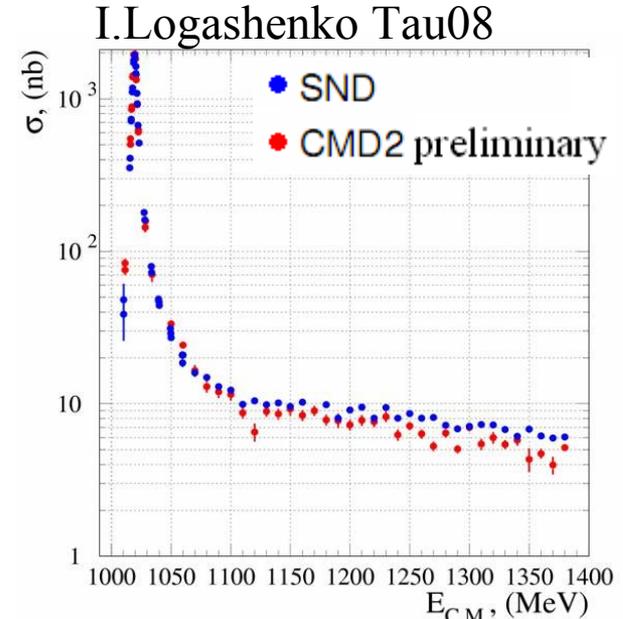
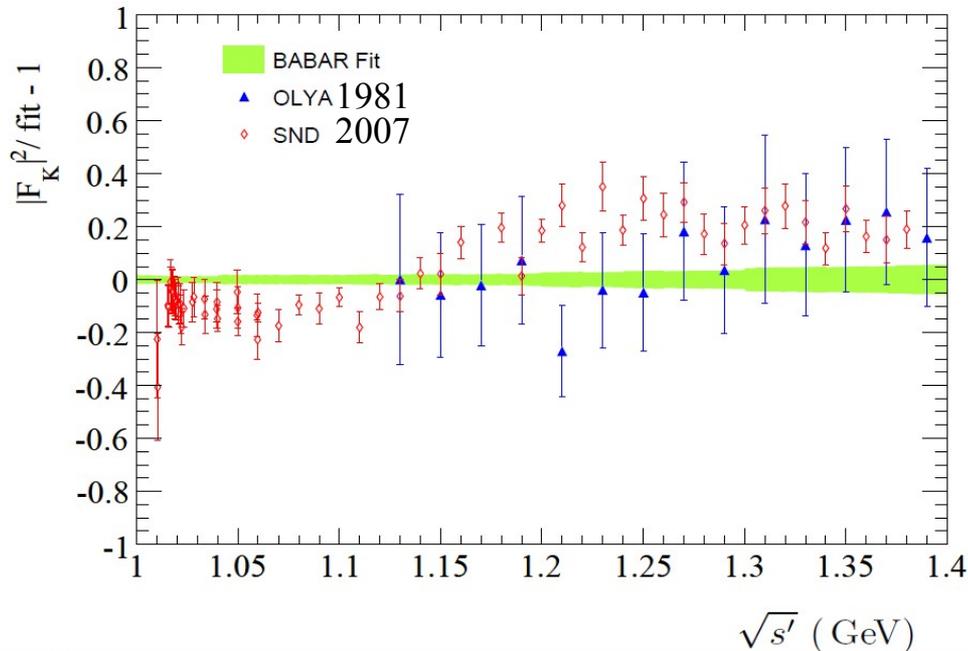
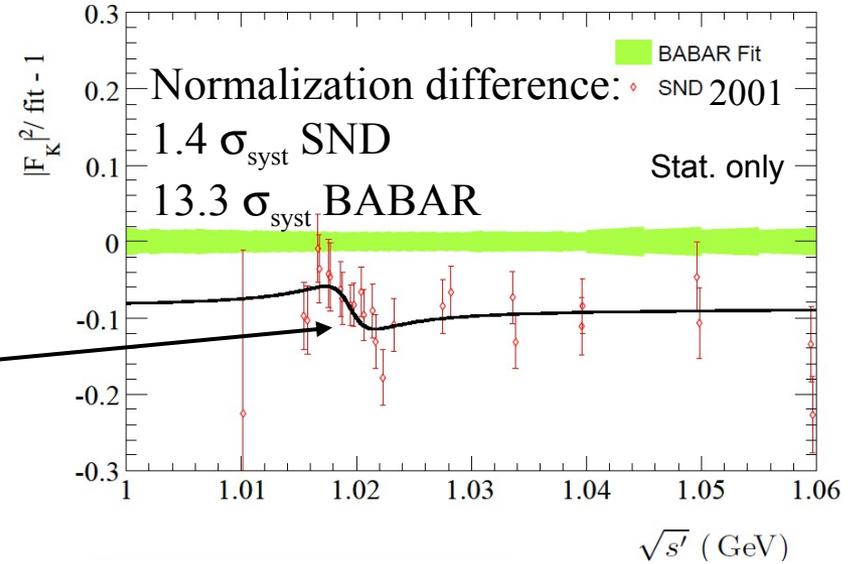
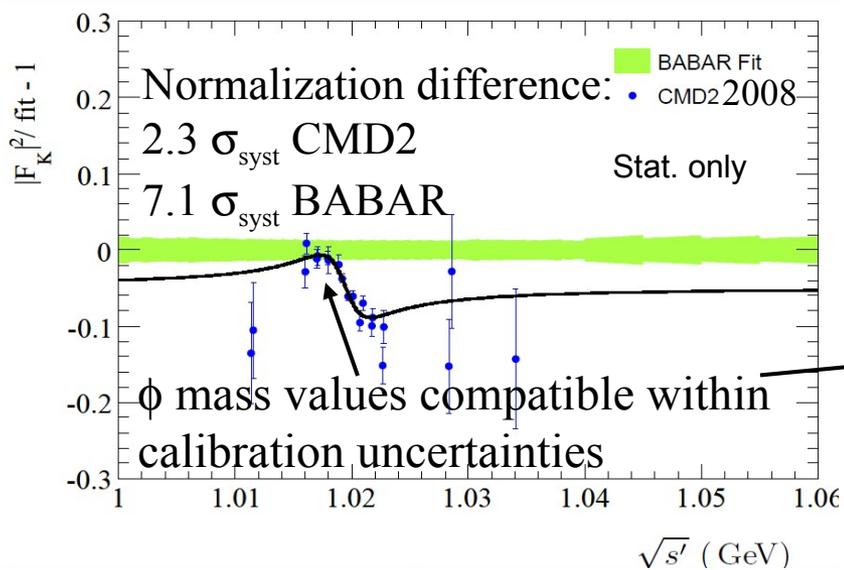
- procedures similar to  $\pi\pi$  analysis
- luminosity  $232 \text{ fb}^{-1}$
- efficiencies obtained from full simulation (AfkQed) and data/MC corrections
  - trigger, tracking, K-ID and mis-ID efficiencies
- background studies, normalization using data, subtraction
- efficiency of the kinematic fit  $\chi^2$  cut
  - additional radiation ISR/FSR
    - studies with muons
  - differences between kaons and muons: secondary interactions, FSR
- unfolding background-subtracted and data/MC corrected mass spectra
- geometrical acceptance and second-order corrections using Phokhara
- ISR effective luminosity from  $\mu\mu\gamma(\gamma)$ : **KK/ $\mu\mu$  ratio**
- mass-dependent systematic uncertainties, best in  $\phi$  region (**0.7%**)
- cross section
- form factor phenomenological fits
- contribution to  $a_\mu$

# Results on the $e^+e^- \rightarrow K^+K^- (\gamma)$ bare cross section with FSR included (small)

→ Use effective ISR luminosity obtained with  $\mu\mu$  sample.



# Comparison to previous experiments



# The $\phi$ parameters

$m_\phi$ ,  $\Gamma_\phi$ , and  $a_\phi$  obtained from the VDM fit ( $\phi$  + other vector mesons) of the form factor

BABAR:

$$m_\phi = 1019.51 \pm 0.02 (\pm 0.05) \text{ MeV}$$

$$\Gamma_\phi = 4.29 \pm 0.04 (\pm 0.07) \text{ MeV}$$

Good agreement with PDG:

$$m_\phi = 1019.455 \pm 0.020 \text{ MeV}$$

$$\Gamma_\phi = 4.26 \pm 0.04 \text{ MeV}$$

From integrated  $\phi$  peak:

$$\Gamma_{ee}^\phi \times B(\phi \rightarrow K^+K^-) = \frac{\alpha^2 \beta^3(s, m_K)}{324} \frac{m_\phi^2}{\Gamma_\phi} a_\phi^2 C_{FS} \text{Final-state correction (Coulomb)}$$

$$\Gamma_\phi^{ee} \times \mathcal{B}(\phi \rightarrow K^+K^-) = (0.6340 \pm 0.0070_{\text{exp}} \pm 0.0037_{\text{fit}} \pm 0.0013_{\text{cal}}) \text{ keV} \quad (1.3\%)$$

$$\text{CMD2 2010: } 0.605 \pm 0.004 \pm 0.013 \text{ keV} \quad (2.2\%)$$

# Charged kaon form factor at large $Q^2$

Predictions based on QCD in asymptotic regime

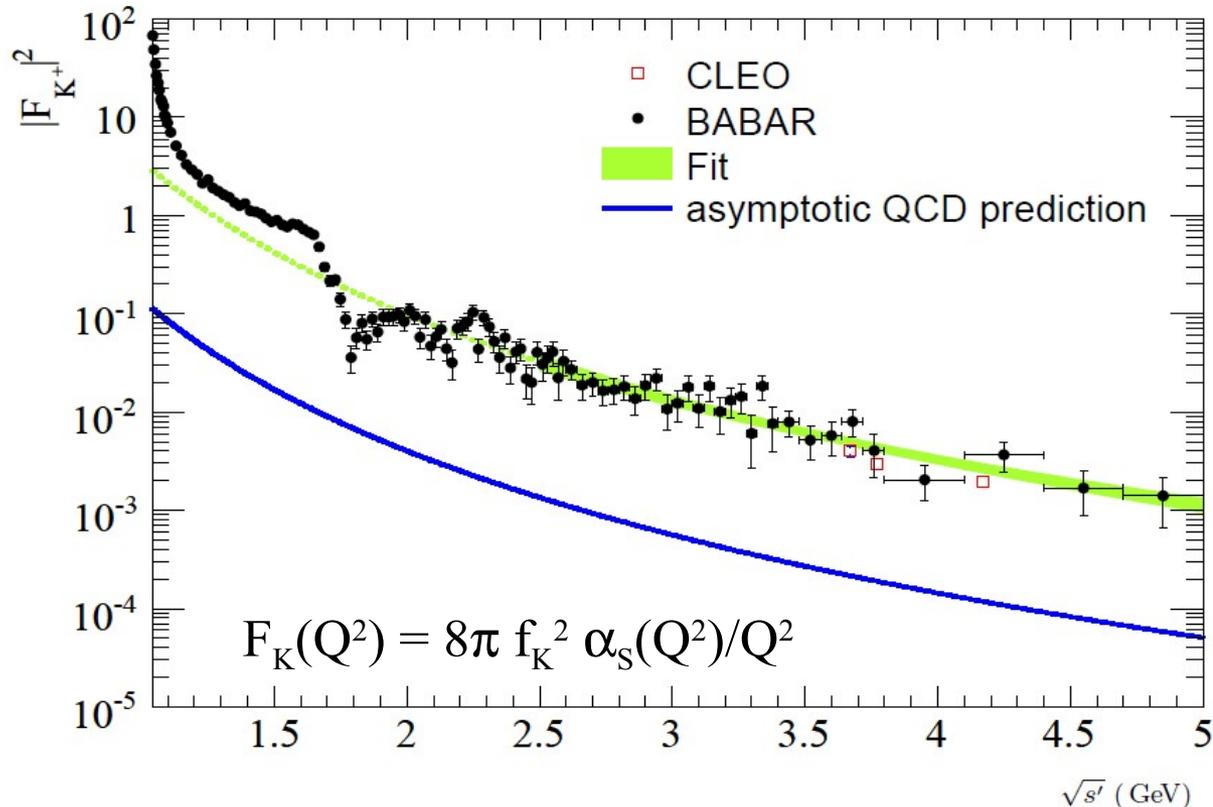
(Chernyak-Zhitnitsky JETP Lett. 1977; Brodsky-Lepage Phys. Lett. B 1979)

→ power law  $F_K \sim \alpha_s(Q^2) Q^{-n}$  with  $n=2$  in good agreement with data  
(2.5-5 GeV  $n=2.04 \pm 0.22$ )

→ but data on  $|F_K|^2$  a **factor  $\sim 20$  above prediction !**

→ no trend in data up to 25 GeV<sup>2</sup> for approaching the asymptotic QCD prediction

→ similar trend observed with  $F_\pi$



# Impact of BABAR data for g-2: $K^+K^-$

BABAR:

$$a_{\mu}^{\text{KK, LO}} [0.98-1.8] \text{ GeV} = (22.93 \pm 0.18 \text{ (stat)} \pm 0.22 \text{ (syst)}) 10^{-10} \quad (1.3\%)$$

DHMZ arXiv:1010.4180 : update of all results before BABAR

$$a_{\mu}^{\text{KK, LO}} [0.98-1.8] \text{ GeV} = (21.63 \pm 0.27 \text{ (stat)} \pm 0.68 \text{ (syst)}) 10^{-10} \quad (3.4\%)$$

BABAR more precise than previous world average by a factor of  $\sim 2.7$

# Conclusions

- Through the ISR method BABAR carried out a complete and consistent program to measure precise cross sections for the dominant channels of  $e^+e^- \rightarrow$  hadrons from threshold to  $\sim 2$  GeV.
- Just a few more channels still in progress.
- BABAR results have a large impact on the knowledge of hadronic vacuum polarization (HVP) contribution to the muon  $g-2$ .
- In addition to HVP there are other applications of these data for QCD tests with finite energy sum rules, complementing similar studies done with hadronic  $\tau$  decays.
- Also (not covered in this talk) BABAR ISR results provide input into hadron spectroscopy, resonance dynamics and measurements of baryon form factors.

# Backup

# $\pi^+\pi^-$ systematic uncertainties

$\sqrt{s}$ ' intervals (GeV)

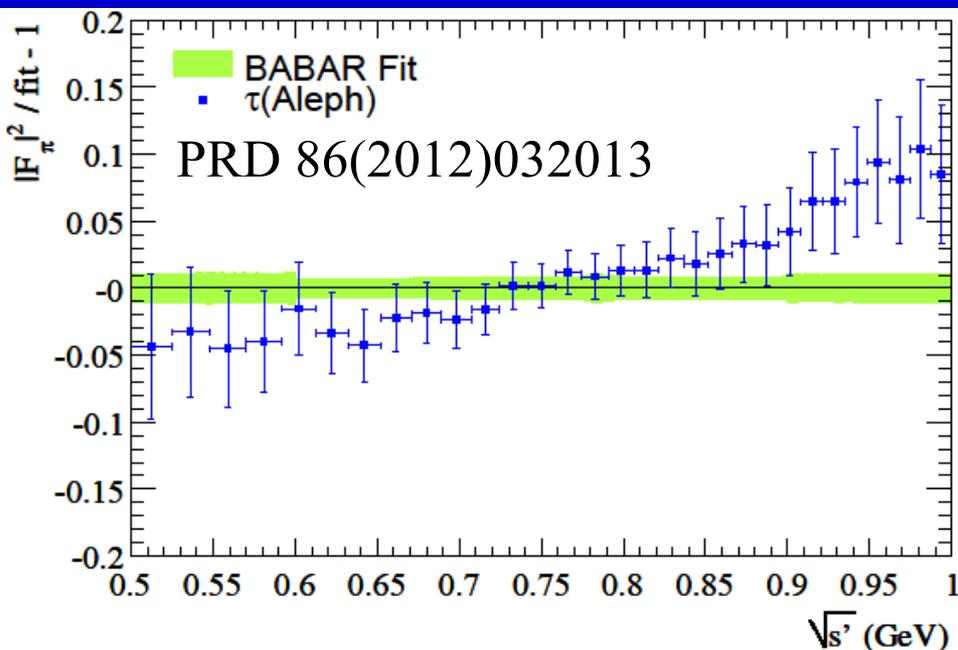
relative uncertainties in  $10^{-3}$

sources	0.3-0.4	0.4-0.5	0.5-0.6	0.6-0.9	0.9-1.2	1.2-1.4	1.4-2.0	2.0-3.0
trigger/ filter	5.3	2.7	1.9	1.0	0.5	0.4	0.3	0.3
tracking	3.8	2.1	2.1	1.1	1.7	3.1	3.1	3.1
$\pi$ -ID	10.1	2.5	6.2	2.4	4.2	10.1	10.1	10.1
background	3.5	4.3	5.2	1.0	3.0	7.0	12.0	50.0
acceptance	1.6	1.6	1.0	1.0	1.6	1.6	1.6	1.6
kinematic fit ( $\chi^2$ )	0.9	0.9	0.3	0.3	0.9	0.9	0.9	0.9
correl $\mu\mu$ ID loss	3.0	2.0	3.0	1.3	2.0	3.0	10.0	10.0
$\pi\pi/\mu\mu$ cancel.	2.7	1.4	1.6	1.1	1.3	2.7	5.1	5.1
unfolding	1.0	2.7	2.7	1.0	1.3	1.0	1.0	1.0
ISR luminosity	3.4	3.4	3.4	3.4	3.4	3.4	3.4	3.4
sum (cross section)	13.8	8.1	10.2	5.0	6.5	13.9	19.8	52.4

$\rho$

Dominated by particle ID ( $\pi$ -ID, correlated  $\mu\mu \rightarrow \pi\pi$ ,  $\mu$ -ID in ISR luminosity)

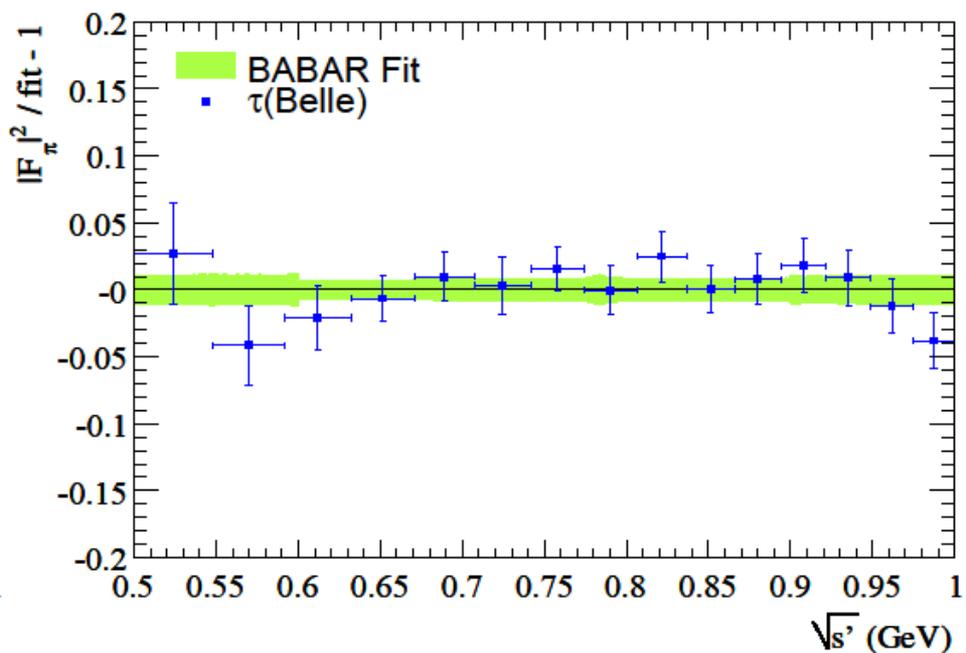
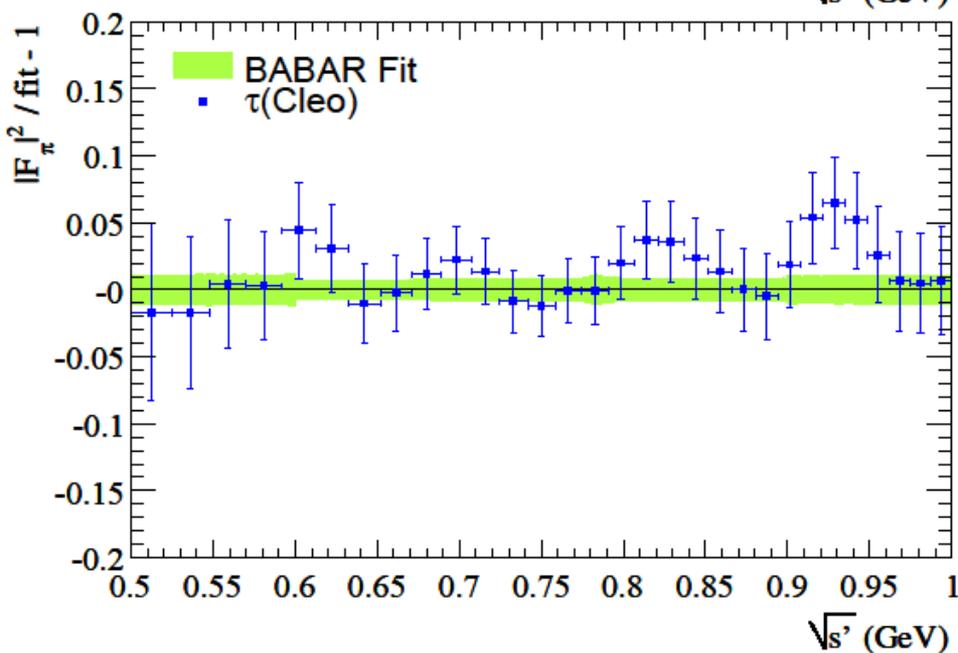
# BABAR vs. IB-corrected $\tau$ data (0.5-1.0 GeV)



Relative comparison w.r.t. BABAR of  $\tau$  spectral functions corrected for isospin-breaking (IB)

IB corrections: radiative corr.,  $\pi$  masses,  $\rho$ - $\omega$  interference,  $\rho$  masses/widths

Each  $\tau$  data normalized to its own BR



# A phenomenological fit to the K form factor

$$F_K(s) = (a_\phi BW_\phi + a_{\phi'} BW_{\phi'} + a_{\phi''} BW_{\phi''})/3$$

$$+ (a_\rho BW_\rho + a_{\rho'} BW_{\rho'} + a_{\rho''} BW_{\rho''} + a_{\rho'''} BW_{\rho'''})/2$$

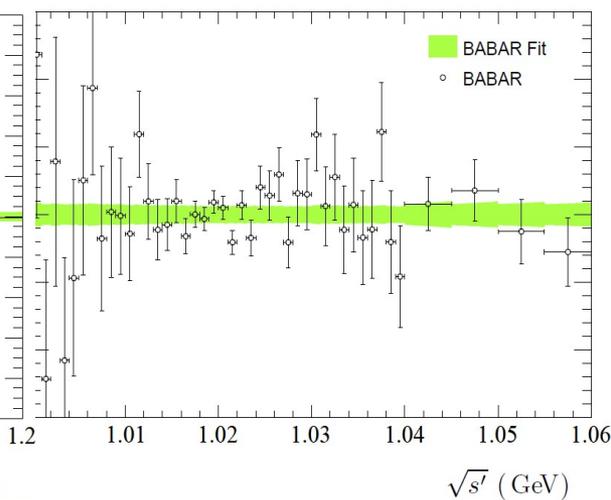
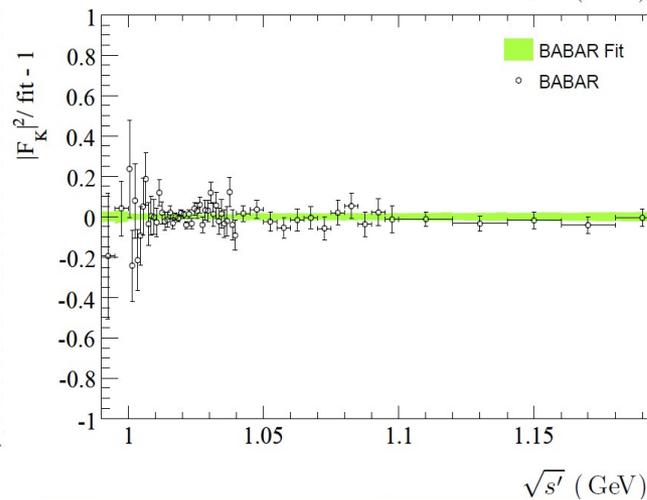
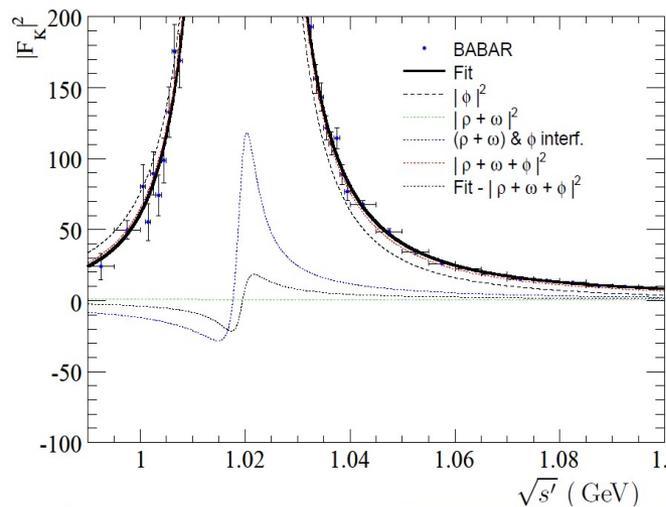
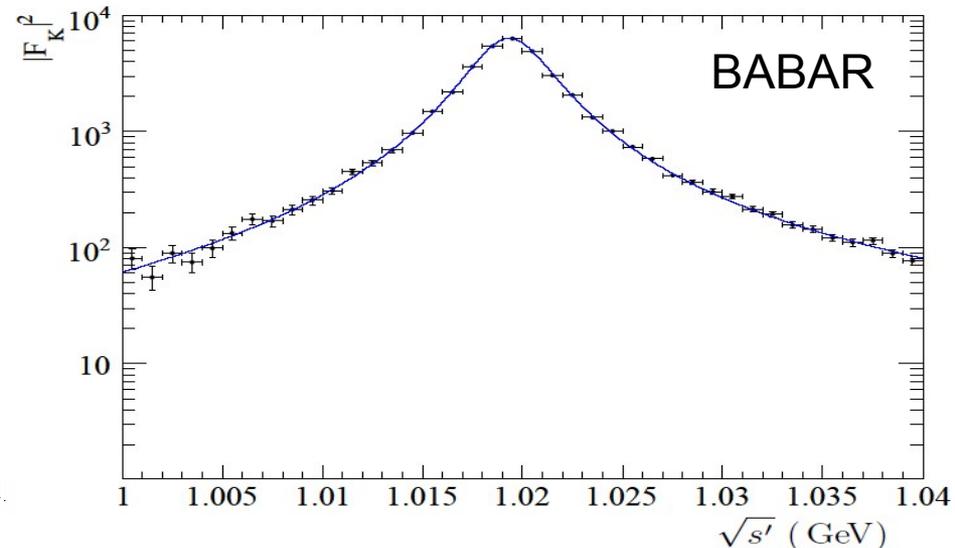
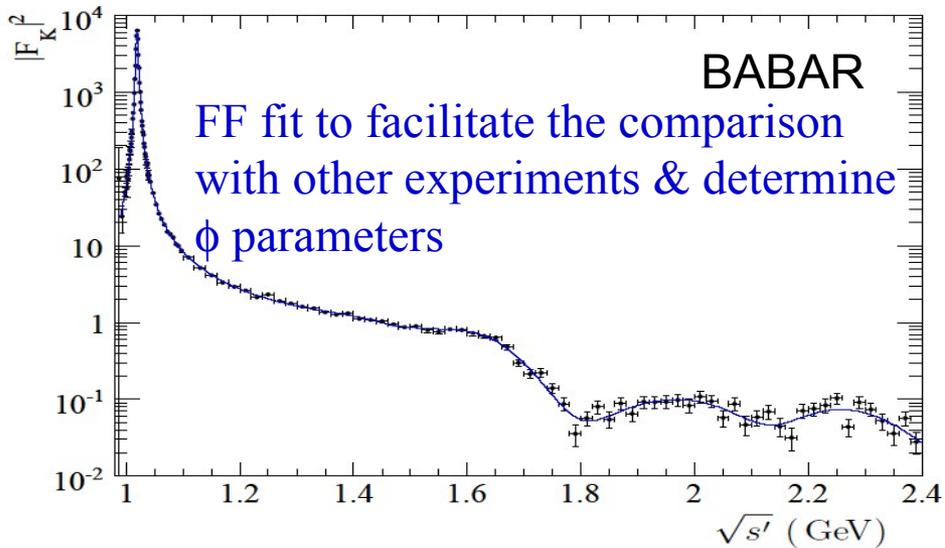
$$+ (a_\omega BW_\omega + a_{\omega'} BW_{\omega'} + a_{\omega''} BW_{\omega''} + a_{\omega'''} BW_{\omega'''})/6$$

Kuehn et al.

$$a_\phi + a_{\phi'} + a_{\phi''} = 1,$$

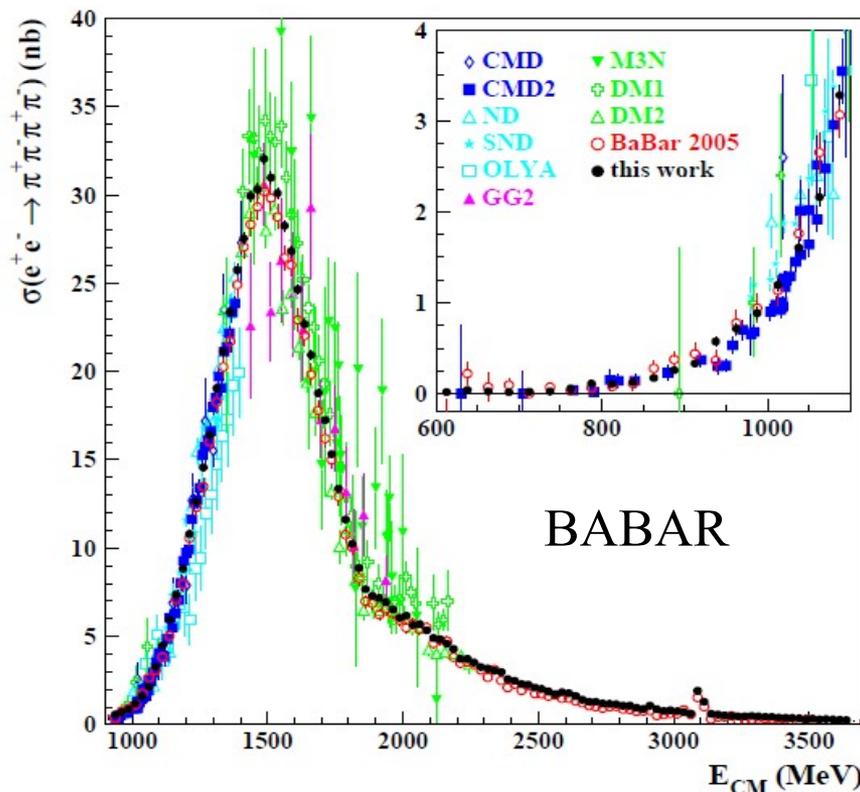
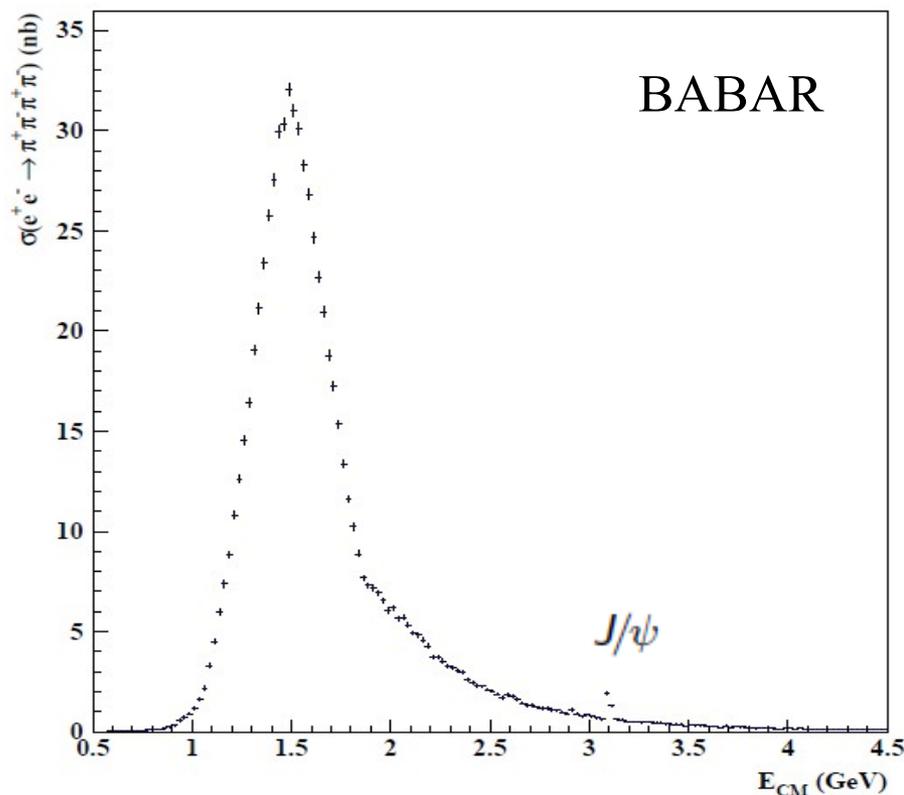
$$a_\rho + a_{\rho'} + a_{\rho''} + a_{\rho'''} = 1,$$

$$a_\omega + a_{\omega'} + a_{\omega''} + a_{\omega'''} = 1.$$



# New results: $e^+e^- \rightarrow \pi^+\pi^-\pi^+\pi^-$

published in 2012 based on the full BABAR statistics (454 fb<sup>-1</sup>)  
previous publication on 89 fb<sup>-1</sup> only



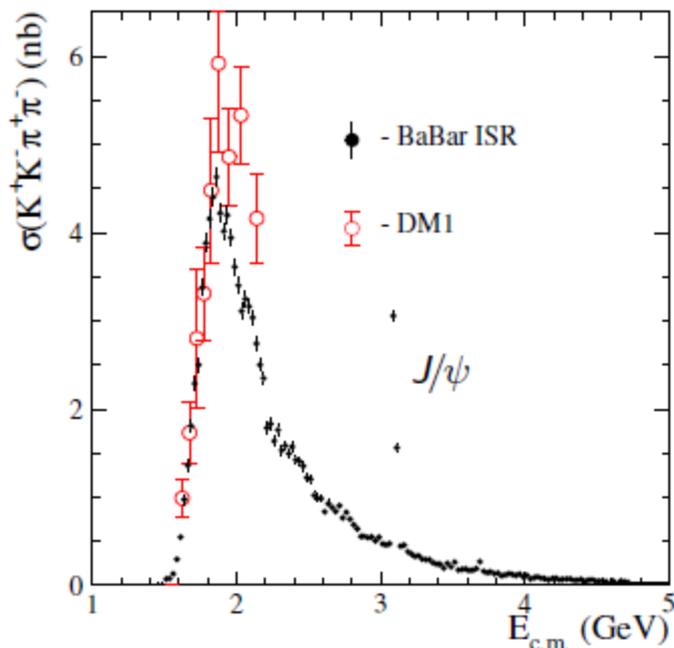
- Systematic uncertainties  
2.4% in peak region (1.1-2.8 GeV)  
11% (0.6-1.1 GeV)  
4% (2.8-4.0 GeV)
- $J/\psi$  visible

- < 1.4 GeV: agreement with previous *BABAR* results, SND and CMD-2 data
- > 1.4 GeV: highest precision (DM2, 20%)

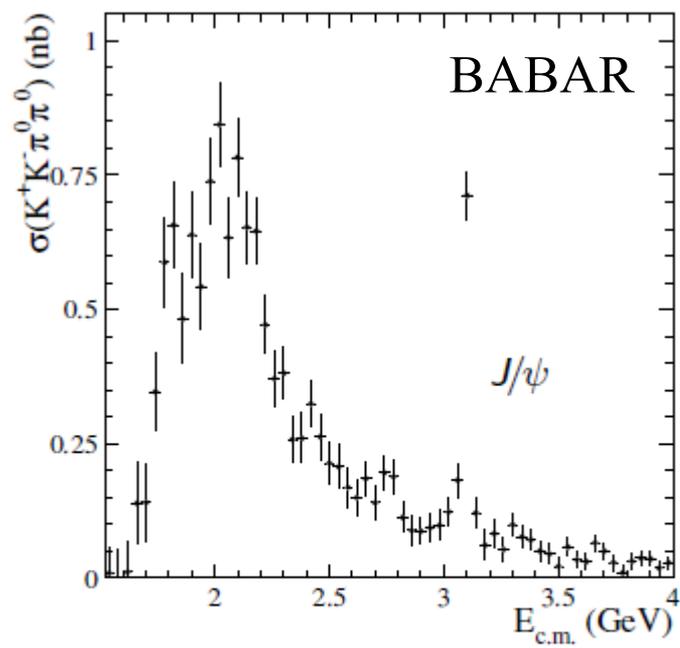
# New results: $e^+e^- \rightarrow K^+ K^- \pi^+ \pi^-$ , $K^+ K^- \pi^0 \pi^0$

Published in 2012 based on the full BABAR statistics (454 fb<sup>-1</sup>)

→ huge improvement compared to existing data



- syst. uncertainty: 4 - 11%
- resolution: 4.2 - 5.5 MeV
- $J/\psi$  clearly visible



- syst. uncertainty: 7 - 16%
- resolution: 8.8 - 11.2 MeV
- $J/\psi$  clearly visible

Cross sections dominated below 1.8 GeV by  $K^*(892)^0 K^+ \pi^-$  and  $K^*(892)^+ \pi^+ \pi^-$   
important to know resonance dynamics to estimate unmeasured final states for  $g-2$  integral

# Impact of BABAR data for g-2: $2(\pi^+ \pi^-)$

BABAR results:

$$a_{\mu}^{4\pi, LO} [0.6-1.8] \text{ GeV} = (13.64 \pm 0.03 \text{ (stat)} \pm 0.36 \text{ (syst)}) 10^{-10} \text{ (2.6\%)}$$

DEHZ 2003: all results but BABAR 2007:

$$a_{\mu}^{4\pi, LO} [0.6-1.8] \text{ GeV} = (13.95 \pm 0.90 \text{ (exp)} \pm 0.23 \text{ (rad*)}) 10^{-10} \text{ (6.7\%)}$$

\* missing radiative corrections

DHMZ 2011: all results but BABAR 2012:

$$a_{\mu}^{4\pi, LO} [0.6-1.8] \text{ GeV} = (13.35 \pm 0.10 \text{ (stat)} \pm 0.52 \text{ (syst)}) 10^{-10} \text{ (4.0\%)}$$

BABAR more precise than previous world average by a factor of 2.6