

Model-independent constraints on the evolution of the dark energy using WMAP and SN-Ia data

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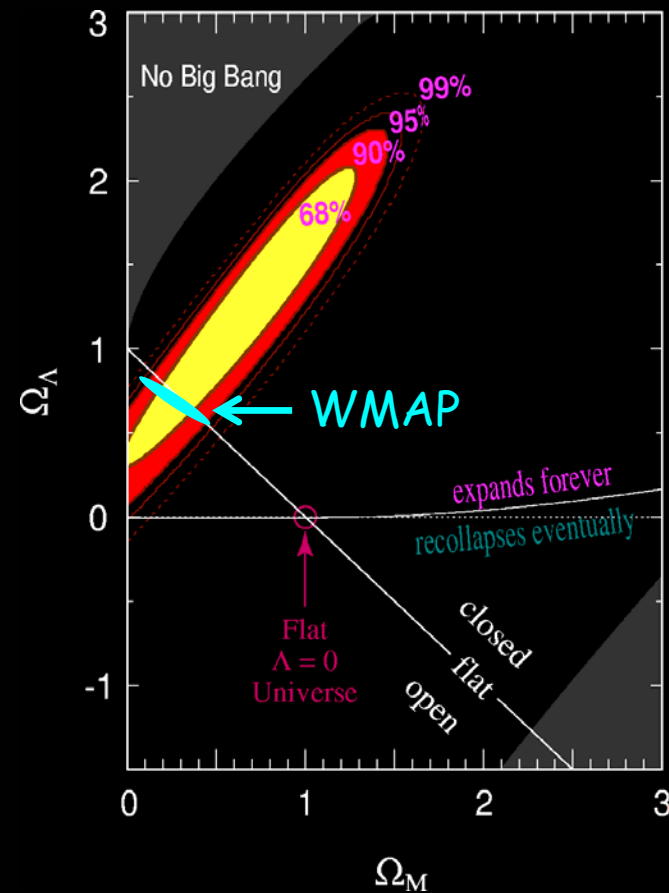
& B.A. Bassett, E.J. Copeland, P.S. Corasaniti,
M. Kunz, J. Silk and C. Ungarelli

Outline

- Why Dark Energy?
- What is it?
- How do we investigate it?
- What did we learn?
- The End

Why do we need Dark Energy?

- Is the Universe low density?
 - Matter + radiation only 30% of critical density
- SN-Ia suggests not the case
- WMAP: $\Omega \approx 1$
- Extra component (Λ) with negative pressure



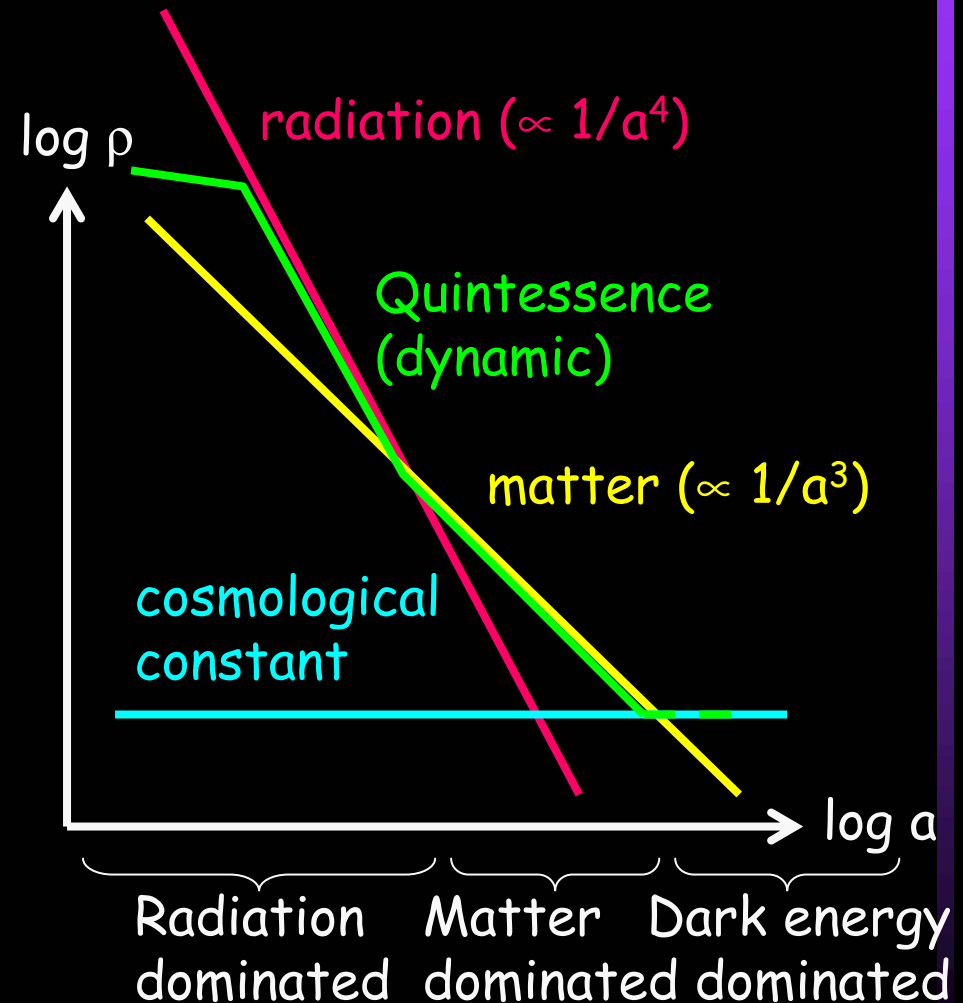
What is it?

Cosmological Constant

- estimate of energy of vacuum using QM gives wrong answer

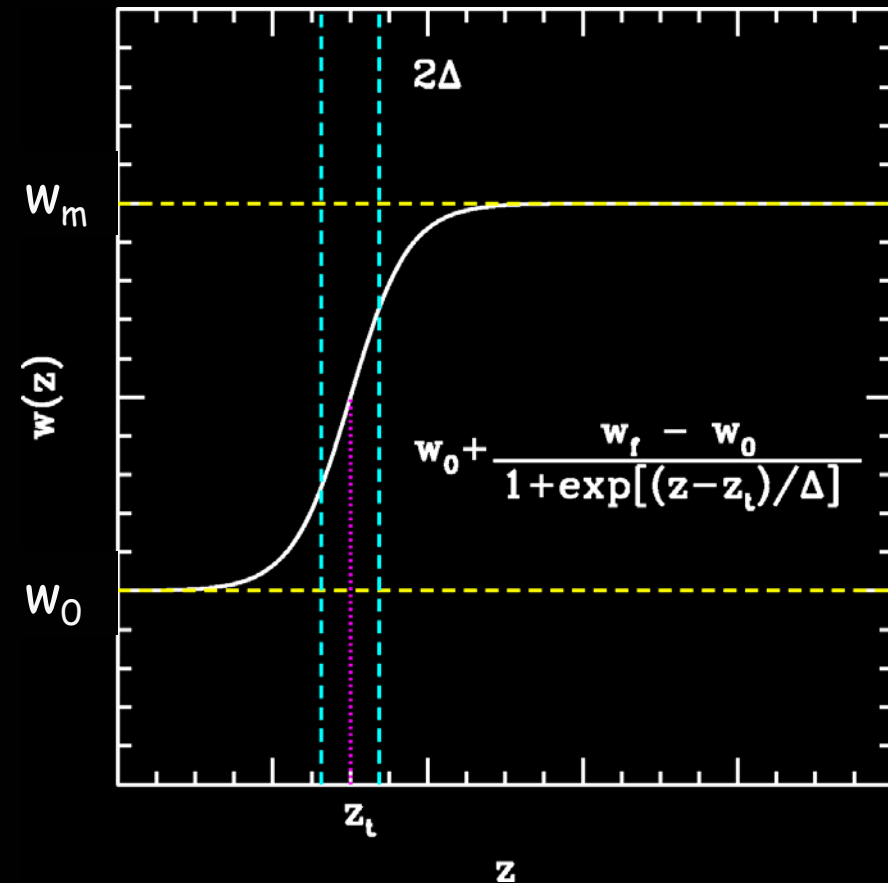
Quintessence

- Can be dynamically linked to background evolution
- Tracker models make i.c.s less important
- Typical potentials include exponentials, inverse power law, exponentials with polynomials, etc



How do we investigate it?

- Impose some formulation for the equation of state
- Transition at given redshift z_{\dagger} with width Δ
- Expect equation of state today, $w_0 < -1/3$ (recent acceleration)
- To avoid fine-tuning i.c.s, $w_m > -1/3$

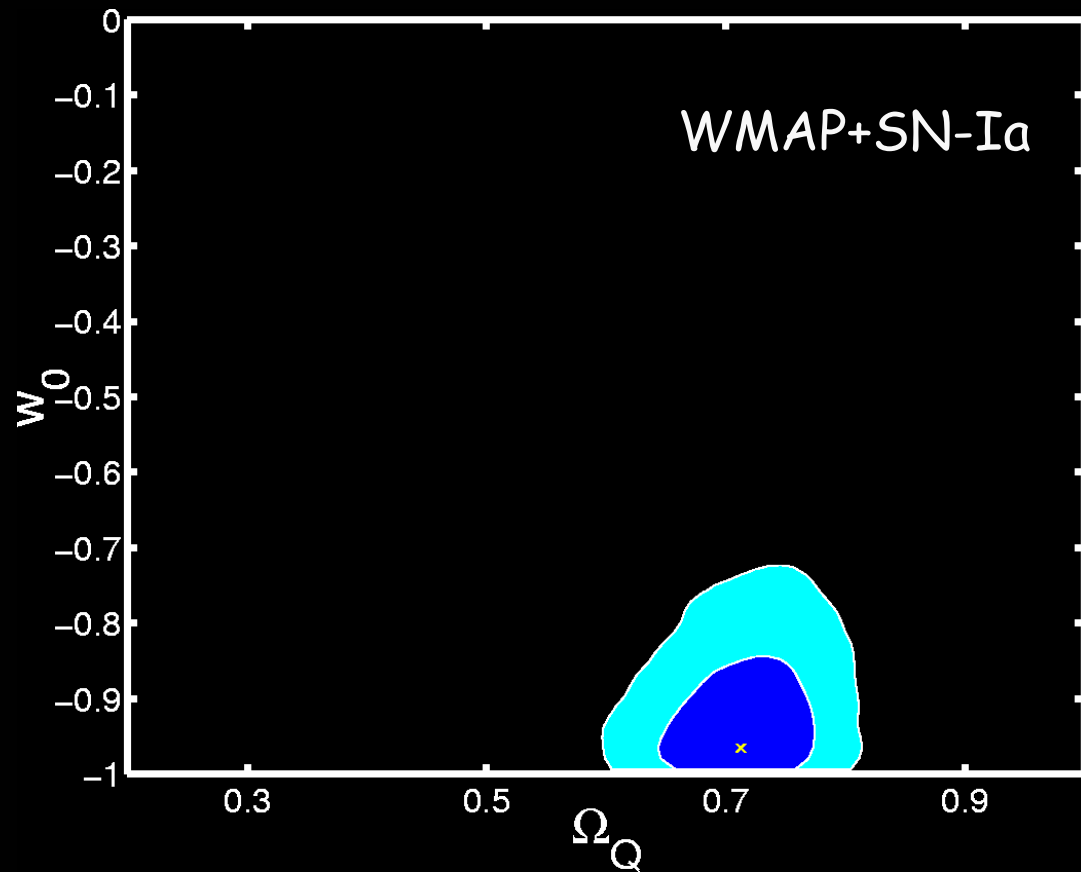


Outline of analysis

- Compute model predictions for different parameters (we vary H_0 , Ω_Q , Ω_b , a_+ , w_0 , w_m , Δ , n_s , τ and the overall normalisation)
- Compare predictions with WMAP + SN-Ia data sets
- Use Markov-Chain Monte Carlo code with modified cmbfast to calculate likelihood surfaces
- Use likelihood data to constrain parameters

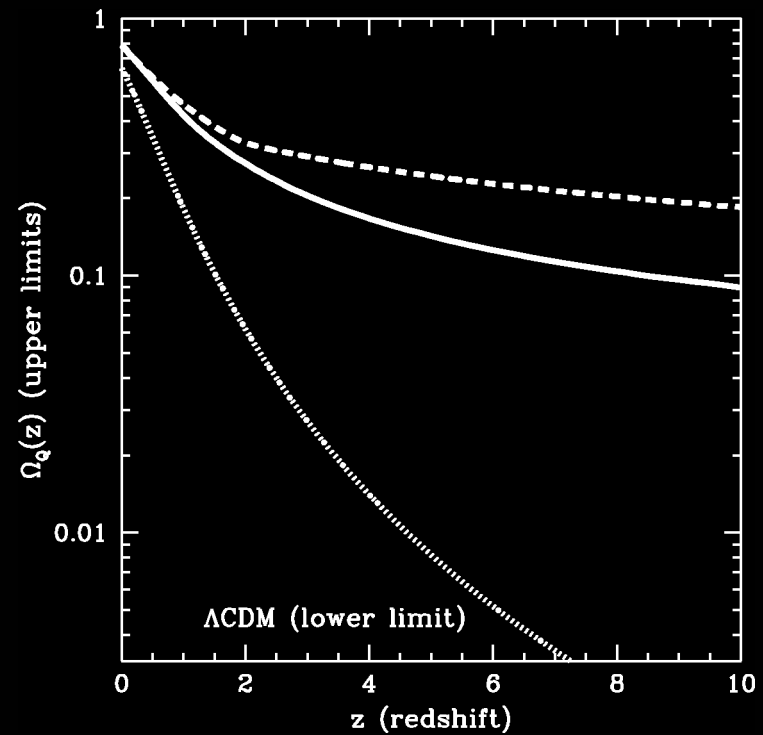
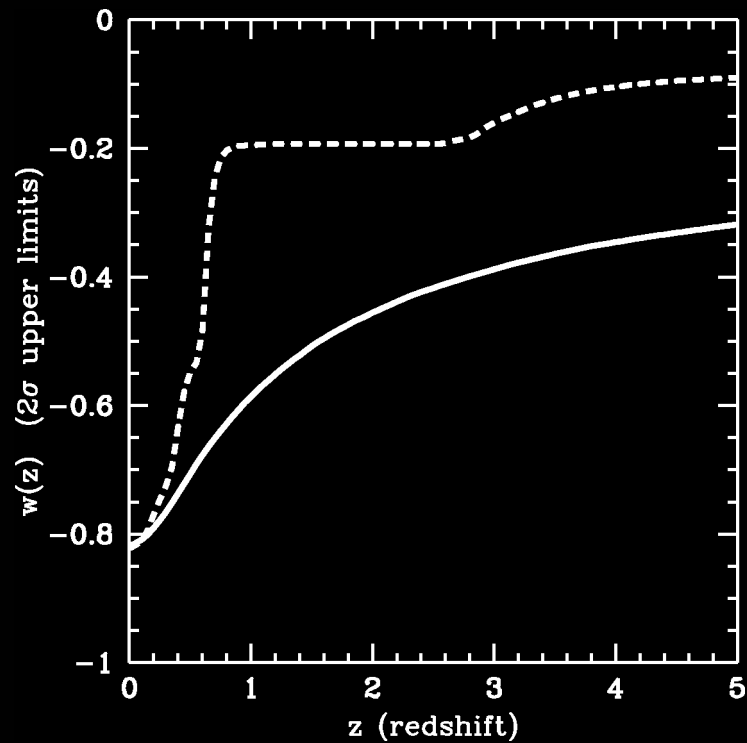
Degeneracies

- Degeneracy in CMB
- Similar in SN-Ia, but runs in other direction
- Combine for tighter constraints.



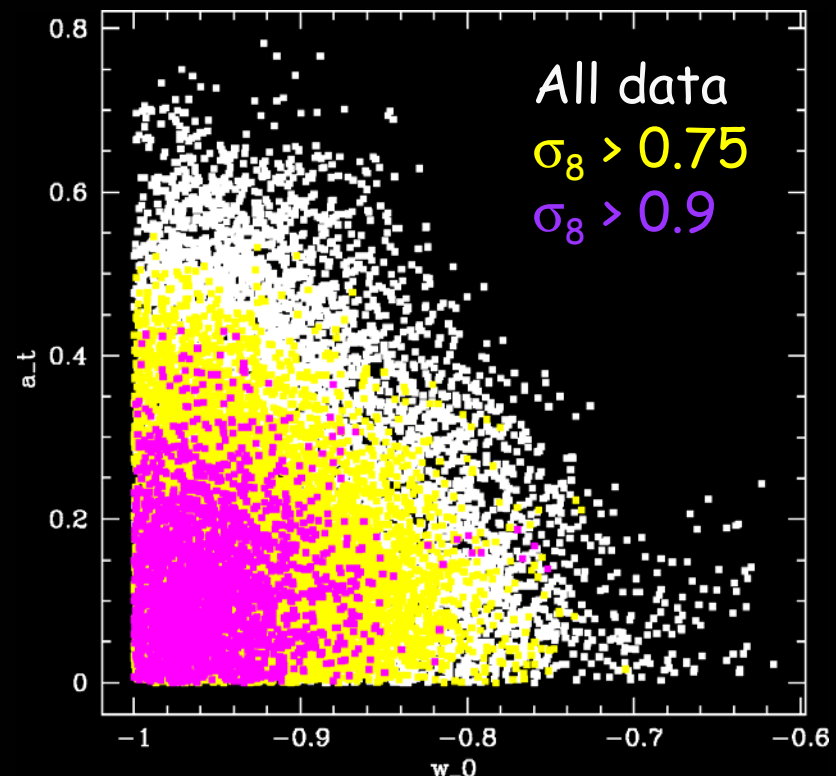
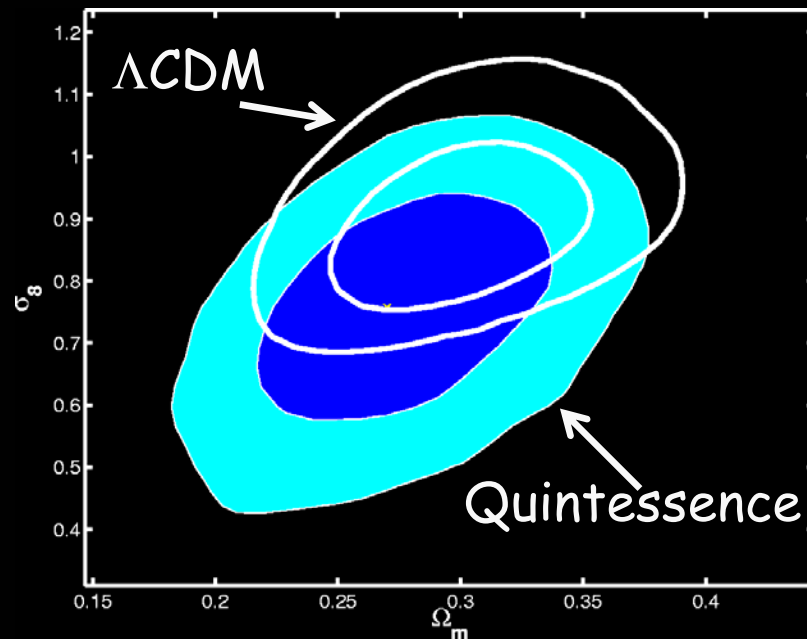
What did we learn?

- Strong constraints only for $z < 1$
- Equation of state w always less than zero



Results (II)

- ISW changes CMB normalisation
- This in turn changes $P(k)$ amplitude and σ_8 (possible to distinguish between Q and Λ)
- But measuring σ_8 is difficult.



Conclusions

- Equation of state today, $w_0 < -0.82$.
- No strong change in w for $z < 1$.
- Can use clustering to distinguish QCDM and Λ CDM
- No significant improvement over Λ CDM

The End

